# REPUBLIC OF TURKEY YILDIZ TECHNICAL UNIVERSITY <br> GRADUATE SCHOOL OF NATURAL AND APPLIED SCIENCES 

## A BRIEF REVIEW OF TRANSFERRING OF TRANSIENT SIGNALS ALONG HOLLOW WAVEGUIDE

SEDEF ŞEMŞIT

MSc. THESIS
DEPARTMENT OF MATHEMATICAL ENGINEERING PROGRAM OF MATHEMATICAL ENGINEERING

ADVISER
ASSOC. PROF. DR. KEVSER KÖKLÜ

# REPUBLIC OF TURKEY <br> <br> YILDIZ TECHNICAL UNIVERSITY <br> <br> YILDIZ TECHNICAL UNIVERSITY <br> GRADUATE SCHOOL OF NATURAL AND APPLIED SCIENCES 

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## Thesis Adviser

Assoc. Prof. Dr. Kevser KÖKLÜ
Yıldız Technical University

Co- Adviser
Assist. Prof. Dr. Emre EROĞLU
Kırklareli University

## Approved By the Examining Committee

Assoc. Prof. Dr.Kevser KÖKLÜ
Yıldız Technical University

Assist. Prof. Dr. Emre EROĞLU
Kırklareli University

Prof. Dr. İsmail KÖMBE
İstanbul Commerce University

Assist. Prof. Dr. Nihan ÇINAR
Kırklareli University

Assist. Prof. Dr. Hasan Hüseyin YURTÇU
Yıldız Technical University

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## LIST OF SYMBOLS

| R | 3-component position vector |
| :---: | :---: |
| $\nabla$ | 3 -component Nabla operator |
| r | 2-component position vector |
| $\nabla_{\perp}$ | Transverse part of $\nabla$ |
| $\partial_{t}$ | Partial derivative operator with respect to time |
| z | Unit vector directed along the Oz -axis |
| n | Unit vector outer normal to domain $S$ |
| 1 | Unit vector tangential to the contour $L$ |
| E | Electric field vector |
| H | Magnetic field vector |
| $\varepsilon_{0}$ | Permitivity of free space |
| $\mu_{0}$ | Permeability of free space |
| c | Speed of light in space |
| $v_{m}$ | Eigenvalues of Neumann problem |
| $\psi_{m}$ | Eigenvectors of Neumann problem |
| $\kappa_{m}$ | Eigenvalues of Dirichlet problem |
| $\phi_{m}$ | Eigenvectors of Dirichlet problem |
| $\mathbf{E}_{m}(\mathbf{R}, t)$ | Transverse part of Electric field vector |
| $\mathbf{H}_{m}(\mathbf{R}, t)$ | Transverse part of Magnetic field vector |
| $h_{m}(z, t)$ | Time-dependent modal amplitudes of magnetic field vector |
| $e_{m}(z, t)$ | Time-dependent modal amplitudes of electric field vector |
| $A_{m}^{h}$ | Normalization constant of magnetic field |
| $A_{m}^{e}$ | Normalization constant of electric field |

# LIST OF ABBREVIATIONS 

EAE Evolutionary Approach to Electromagnetics
FDTD Finite Difference Time Domain
KGE Klein-Gordon Equation
TE Transverse Electric
TM Transver Magnetic

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# ABSTRACT 

# A BRIEF REVIEW OF TRANSFERRING OF TRANSIENT SIGNALS ALONG HOLLOW WAVEGUIDE 

Sedef ŞEMŞİT<br>Department of Mathematical Engineering<br>MSc. Thesis<br>Adviser: Assoc. Prof. Dr. Kevser KÖKLÜ<br>Co-adviser: Assist. Prof. Dr. Emre EROĞLU

In this thesis, the purpose is to bring together the studies done before. Evolutionary Approach to Electromagnetics Theory (EAE), which is an anlytical method and gives the solutions in directly time-domain, is aimed to use. Two main problems are introduced to analyze transferring of transient signals along hollow waveguide. First one, Modal Basis Problem corresponds to Neumann and Dirichlet boundary-eigenvalue problems. These problems gives us eigenfuctions corresponds to eigenvalues. Second one is Modal Amplitude Problem which derived from Klein-Gordon equations. As a result of this, transverse electric (TE) and transverse magnetic (TM) modes are obtained in time-domain. Each component of these modal fields is expressed as a product of modal base element which is a vector function of transverse coordinates and a scalar modal amplitude that depends on time $t$ and axial coordinate $z$. Also, in this study, Bessel and Airy functions are analyzed in detail and they are used to expressed Modal amplitudes.

Key words: Maxwell's equation, Neumann and Dirichlet boundary-eigenvalue problems, Klein-Gordon equation, Airy functions, Bessel functions

## ÖZET

# BOŞ DALGA KLAVUZU BOYUNCA GEÇíCİ SİNYALLERİN İLETILMESİNIN KISA BİR İNCELEMESİ 

Sedef ŞEMŞiT<br>Matematik Mühendisliği Anabilim Dalı<br>Yüksek Lisans Tezi<br>Tez Danışmani: Doç. Dr. Kevser KÖKLÜ<br>Eş Danı̧̧man: Yrd. Doç. Dr. Emre EROĞLU

Bu tezde amaç daha önce yapılmış çalışmaları bir araya getirmektir. Analitik bir metod olan ve çözümleri direk olarak zaman-domeininde veren Elektromanyetik Teoriye Evrimsel Yaklaşım (ETEY) metodunu kullanmak amaçlanmıştır. Dalga kılavuzu boyunca geçici sinyallerin iletilmesini incelemek için iki temel problem tanıtılmıştır. Birincisi, Neumann ve Dirichlet sınır-özdeğer problemlerine karşılık gelen Modal Baz Problemidir. Bu problemlerden özdeğerlere bağlı özfonksiyonlar elde edilmiștir. İkincisi ise Klein-Gordon denkleminden elde edilen Modal Genlik Problemidir. Bunların sonucunda enlemsel elektrik (TE) ve enlemsel manyetik modları zamandomeininde elde edilmiştir. Bu modal alanların her bir elemanı, boylamsal koordinatların vektör fonksiyonu olan modal baz ve $t$ zaman ve $z$ eksenel koordinatına bağlı olan modal genlik çarpımıyla ifade edilmişlerdir. Ayrıca bu çalışmada Bessel ve Airy fonksiyonları deletaylı olarak incelenmiş ve modal genlikler bu fonksiyonlar yardımıyla bulunmuştur.

Anahtar kelimer: Maxwell denklemleri, Neumann ve Dirichlet sinır-özdeğer problemleri, Klein-Gordon denklemi, Airy fonksiyonları, Bessel fonksiyonları.

## CHAPTER 1

## INTRODUCTION

### 1.1 Literature Review

In Electromagnetics, there are three canonical boundary value problem with given initial conditions for the electromagnetic field sought, namely: Cavity Problem, Waveguide Problem, and External Problem. In this thesis, the waveguide problem which is conanical problem is studied in a subdomain of Euclidean space.

The time-domain studies in electromagnetic field theory can be investigated in two groups, in general. The first one is numerical methods; Finite Difference Time Domain (FDTD) is the most popular one among these methods. Efficiency of the method is abundance of data which has obtained during oscillation. Compariable process is permitted to reach exact results via rich data. The second one is analytical methods that divided into two subgroups. The first one consists of frequency domain usually performed concerned with Fourier integral transform or Laplace transform. The solutions obtained from the Fourier integral transform is needed to find time-domain solutions in the final step. A new alternative (to the time-harmonic field concept) approach in studying the time-domain modes was developed within the framework of four-dimensional relativistic formalism in electrodynamics.

Eventually, one more alternative approach called Evolutionary Approach to Electromagnetics (EAE) was suggested in 80s. It is destined for the time-domain theory of both the cavity and waveguide modes. The theory is based on evolution equations was kept $\partial t$ in time-domain of Maxwell's equations and Neumann-Dirichlet boundaryvalue problems, analytically. In the method, electromagnetic fields are solved by extracting Maxwell's equation system in time-domain. The solution is attained by observing along waveguide

The EAE method is based on the solution of sequential two autonomous problems, basically. The first one is a "modal basis problem" corresponds to well-studied Dirichlet and Neumann boundary eigenvalue problems. This involves two complete sets as the Transverse-Electric (TE) and Transverse-Magnetic (TM) modes. Because of the fact that the generating scalar potentials for the TE and TM modes are actually the same as the time-harmonic modes, one can freely use the methods, which have been developed in the frequency domain, and use as well even ready results obtained for the complex waveguide configurations. The second one is a "time-dependent modal amplitude problem" corresponds to Klein-Gordon Equation (KGE) with the axial coordinate and time. Main effort of theory is addressed to obtain the analytical solution to the KGE leading to the timedependent modal amplitudes unlike to those in the time-harmonic waves .

### 1.2 Objective of the Thesis

In this thesis, the purpose is to solve Maxwell's equation system analytically via Laplacian and evolution equation that preserve $\partial_{t}$ in time-domain. Electromagnetic fields will be separated from Maxwell's equation system and solved in time-domain. The solution is obtained by investigating along the waveguide. This study is composed as follows.

In Section II, the above mentioned executive summary is presented, definition about the waveguide is given and time-domain problems are solved

In Section III, the general properties of the waveguide time-domain modes are considered, namely. Completeness of the modal sets is discussed.. Conservation of energy law is given for every time-domain mode. The initial conditions for the modal amplitudes are mentioned. The causality principle is imposed on the time-domain modes.

In Section IV, a particular case of Bessel and Airy functions are considered in detail and time-dependent modal amplitudes, which are expressible via Bessel and Airy functions with time in their arguments, are obtained.

In Section V, graphical results are presented by using Mapple programme.

### 1.3 Hypothesis

There are lots of study about EAE method. But all these studies fouced on specific titles. Therefore, it is necessary to gather them in one study. So this is the major contribution of this thesis.

As understood from the title of the thesis, we are going to reconsider all the works have been done up to this time. The works are stated below:

- Serkan Aksoy (Aksoy [1]) gave general information about the problem like basis formed in wave boundary operator, projection of vector fields and Maxwell's equation on the basis and numerical applications for rectangular waveguide
- Özlem Akgün (Akgün [2]) found the solution of Klein-Gordon Equation (modal amplitudes) by using Miller's fifth case as a product of Airy function and she investigated relativistic effects on the wave propagation in a hollow waveguide
- Emre Eroğlu (Eroğlu [3]) solved KGE by using Miller's second case as a product of Bessel function, gave detailed solution of Bessel differential equation, studied spherical and cylindrical Bessel function and he also focused on surplus of energy and energy.
- Özlem Işık (Işık [4]) analyzed wave boundary operator and surplus of enery and expressed the modal amplitudes in terms of Airy functions.
- Nevra Eren (Eren [5]) investigated surplus of enery and airy function in detail.


## CHAPTER 2

## THE PROBLEM of TIME-DOMAIN MODES

### 2.1 The Waveguide

The waveguide is a device that transfer electromagnetic energy efficiently which means with minimum extinction and degradation. Propagation of electromagnetic fields can be classfied as guided and unguided. The waves in guided propagation are transmitted from one point to another by following a particular path. The waves in unguided propagation spread or beam on an open space. There are three types of waveguide namely; circular, elliptic and rectangular.

We consider a hollow rectangular waveguide (as in Figure 2.1) directed along Oz -axis and bounded by conductive metal walls that means the waveguide is perfect electric conductor.


Figure 2.1 The rectangular waveguide along Oz-axis
where $\varepsilon_{0}$ is dielectric constant for free space (i.e., permitivity) and $\mu_{0}$ is magnetic constant for free space (i.e., permeability).

### 2.2 Formulation of the Problem

We consider a hollow waveguide having domain $S$ directed along $O z$-axis. The domain $S$ is bounded by closed singly-connected contour $L$. We also assume that the contour has an arbitrary but smooth enough shape which means any possible inner angles (measured within $S$ ) can not exceed $\pi$. In the rest of the study, mutually orthogonal unit vectors ( $\mathbf{z}, \mathbf{l}, \mathbf{n}$ ) were used where $\mathbf{z} \times \mathbf{l}=\mathbf{n}, \mathbf{l} \times \mathbf{n}=\mathbf{z}$ and $\mathbf{n} \times \mathbf{z}=\mathbf{l}$. The unit vector $\mathbf{n}$ is outward normal to domain $S$, the unit vectors $\mathbf{z}$ and $\mathbf{l}$ are tangential to the $O z$-axis and contour $L$, respectively.

In this study, 3 -component position vector $\mathbf{R}$ and the operator $\nabla$ are decomposed onto their transverse and longitudinal parts as

$$
\begin{equation*}
\mathbf{R}=\mathbf{r}+\mathbf{z z}, \quad \nabla=\nabla_{\perp}+\mathbf{z} \partial_{z} \tag{2.1}
\end{equation*}
$$

where $\mathbf{z}$ is a unit vector directed along $O z$-axis, $\mathbf{r}$ is 2 -component position vector in the waveguide cross-section $S$ (and also projection of $\mathbf{R}$ on domain $S$ ) and $\nabla_{\perp}$ is the transverse part of $\nabla$. The differential operator $\nabla_{\perp}$ transacts only at the coordinates ( $\mathbf{r}$ ) of transver waveguide.

3 -component electromagnetic field strength vectors $\mathbf{E}_{m}$ and $\mathbf{H}_{m}$ are introduced

$$
\begin{align*}
& \mathbf{E}_{m}(\mathbf{R}, t)=\mathbf{E}(\mathbf{r}, z, t)+\mathbf{z} E_{z}(\mathbf{r}, z, t) \\
& \mathbf{H}_{m}(\mathbf{R}, t)=\mathbf{H}(\mathbf{r}, z, t)+\mathbf{z} H_{z}(\mathbf{r}, z, t) \tag{2.2}
\end{align*}
$$

and the boundary conditions hold over the waveguide surface are

$$
\begin{equation*}
\left.\left(\mathbf{n} \cdot \mathbf{H}_{m}\right)\right|_{L}=0,\left.\quad\left(\mathbf{l} \cdot \mathbf{E}_{m}\right)\right|_{L}=0,\left.\quad\left(\mathbf{z} \cdot \mathbf{E}_{m}\right)\right|_{L}=0 \tag{2.3}
\end{equation*}
$$

where $\mathbf{n}$ is unit vector outer normal to domain $S, \mathbf{z}$ and $\mathbf{l}$ are the unit vectors tangential to the $O z-$ axis and contour $L$, respectively.

### 2.3 TE Time-Domain Modes

To obtain components of TE time-domain modes let's solve Neumann boundary eigenvalue problem for $\nabla_{\perp}$
$\left(\nabla_{\perp}^{2}+v_{m}^{2}\right) \psi_{m}(\mathbf{r})=0$
$\left.\frac{\partial \psi_{m}(\mathbf{r})}{\partial n}\right|_{L}=0$
$\frac{v_{m}^{2}}{S} \int_{S}\left|\psi_{m}(\mathbf{r})\right|^{2} d S=1 N$
where $\partial_{n}=\mathbf{n} \cdot \nabla_{\perp}$ is normal derivative on $L-$ contour, $v_{m}^{2}>0$ are eigenvalues for $m=$ $1,2, \ldots$, the index $m$ is numerical values sorted by increasing order in the real axis and $\psi_{m}(\mathbf{r})$ 's are eigenvectors corresponding to the eigenvalues. In equation (2.4) the first equation is Helmholtz equation, the second one is boundary condition and the third one is normalization condition. The force dimension in equation (2.4) has to be $N$ (Newton) so that physical dimensions, $V_{m}^{-1}$ (Volt per meter) and $A_{m}^{-1}$ (Ampere per meter), are provided for the field vector extensions $\mathbf{E}_{m}$ and $\mathbf{H}_{m}$, respectively.

Example 2.1 Let us consider $L$-contour that is a rectangular waveguide with boundaries $0 \leq x \leq a$ and $0 \leq y \leq b$.

Since

$$
\nabla_{\perp}=\left(\frac{\partial}{\partial x}, \frac{\partial}{\partial y}\right)
$$

we can write Helmholtz equation as

$$
\left(\frac{\partial^{2}}{\partial x^{2}}+\frac{\partial^{2}}{\partial y^{2}}\right) \psi_{m}+v_{m}^{2} \psi_{m}=0 \quad \Rightarrow \quad \frac{\partial^{2} \psi_{m}}{\partial x^{2}}+\frac{\partial^{2} \psi_{m}}{\partial y^{2}}+v_{m}^{2} \psi_{m}=0
$$

By using the separation of variables method, choose $\psi_{m}(\mathbf{r})=X(x) Y(y)$ and substitute it into above equation yields
$\frac{\partial^{2}[X(x) Y(y)]}{\partial x^{2}}+\frac{\partial^{2}[X(x) Y(y)]}{\partial y^{2}}+v_{m}^{2} X(x) Y(y)=0$
or equavalently
$Y \frac{\partial^{2} X}{\partial x^{2}}+X \frac{\partial^{2} Y}{\partial y^{2}}+v_{m}^{2} X Y=0$.

Multiply both sides with $1 / X Y$ gives
$\underbrace{\frac{1}{X} \frac{\partial^{2} X}{\partial x^{2}}}_{-v_{x}^{2}}+\underbrace{\frac{1}{Y} \frac{\partial^{2} Y}{\partial y^{2}}}_{-v_{y}^{2}}+v_{m}^{2}=0 \quad \Rightarrow \quad v_{m}^{2}=v_{x}^{2}+v_{y}^{2}$
The eigenvalues $v_{x}^{2}$ and $v_{y}^{2}$ are choosen as negative because if we choose them as positive the solution of Neumann problem becomes $\psi_{m}(\mathbf{r})=\exp ( \pm i v x) \exp ( \pm i v y)$ which is a divergent function. So, it is not suitable for the EAE method.

Therefore, the solution of above equation can be written as
$X(x)=A \cos \left(v_{x} x\right)+B \sin \left(v_{x} x\right)$
$Y(y)=C \cos \left(v_{y} y\right)+D \sin \left(v_{y} y\right)$
or
$\psi_{m}(\mathbf{r})=X(x) Y(y)=\left[A \cos \left(v_{x} x\right)+B \sin \left(v_{x} x\right)\right]\left[C \cos \left(v_{y} y\right)+D \sin \left(v_{y} y\right)\right]$.

By using boundary condition, the constants of $\psi_{m}(\mathbf{r})$ can be found as stated below.
(i) $0 \leq x \leq a,\left.\frac{\partial \psi_{m}(\mathbf{r})}{\partial y}\right|_{y=0}=0$

$$
\left.\frac{\partial \psi_{m}}{\partial y}\right|_{y=0}=\left.\left[A \cos \left(v_{x} x\right)+B \sin \left(v_{x} x\right)\right]\left[-C v_{y} \sin \left(v_{y} y\right)+D v_{y} \cos \left(v_{y} y\right)\right]\right|_{y=0}=0
$$

which is equivalent to
$\left[A \cos \left(v_{x} x\right)+B \sin \left(v_{x} x\right)\right]\left[-C v_{y} \sin (0)+D v_{y} \cos (0)\right]=0$
or
$\left[A \cos \left(v_{x} x\right)+B \sin \left(v_{x} x\right)\right]\left[0+D v_{y} 1\right]=0$

In this equation $A \cos \left(v_{x} x\right)+B \sin \left(v_{x} x\right) \neq 0$ so
$0+D v_{y} 1=0 \quad \Rightarrow \quad D=0$
(ii) $0 \leq x \leq a,\left.\frac{\partial \psi_{m}(\mathbf{r})}{\partial y}\right|_{y=b}=0$
$\left.\frac{\partial \psi_{m}}{\partial y}\right|_{y=b}=\left.\left[A \cos \left(v_{x} x\right)+B \sin \left(v_{x} x\right)\right]\left[-C v_{y} \sin \left(v_{y} y\right)+D v_{y} \cos \left(v_{y} y\right)\right]\right|_{y=b}=0$
which is equivalent to
$\left[A \cos \left(v_{x} x\right)+B \sin \left(v_{x} x\right)\right]\left[-C v_{y} \sin \left(v_{y} b\right)\right]=0$
Again in this equation $A \cos \left(v_{x} x\right)+B \sin \left(v_{x} x\right) \neq 0$ so

$$
\begin{aligned}
{\left[-C v_{y} \sin \left(v_{y} b\right)\right] } & =0=\sin (0+p \pi), \quad p=0,1,2, \ldots \\
& \Rightarrow v_{y}=\frac{p \pi}{b}
\end{aligned}
$$

(iii) $0 \leq y \leq b,\left.\frac{\partial \psi_{m}(\mathbf{r})}{\partial x}\right|_{x=0}=0$
$\left.\frac{\partial \psi_{m}}{\partial x}\right|_{x=0}=\left.\left[-A v_{x} \sin \left(v_{x} x\right)+B v_{x} \cos \left(v_{x} x\right)\right]\left[C \cos \left(v_{y} y\right)+0\right]\right|_{x=0}=0$
which is equivalent to

$$
\left[-A v_{x} \sin (0)+B v_{x} \cos (0)\right]\left[C \cos \left(v_{y} y\right)\right]=0
$$

or
$\left[0+B v_{x} 1\right]\left[C \cos \left(v_{y} y\right)\right]=0$

In this equation $C \cos \left(v_{y} y\right) \neq 0$ so
$0+B v_{x} 1=0 \quad \Rightarrow \quad B=0$
(iv) $0 \leq y \leq b,\left.\frac{\partial \psi_{m}(\mathbf{r})}{\partial x}\right|_{x=a}=0$
$\left.\frac{\partial \psi_{m}}{\partial x}\right|_{x=a}=\left.\left[-A v_{x} \sin \left(v_{x} x\right)+0\right]\left[C \cos \left(v_{y} y\right)+0\right]\right|_{x=a}=0$
which is equivalent to
$\left[-A v_{x} \sin \left(v_{x} a\right)\right]\left[C \cos \left(v_{y} y\right)\right]=0$

Again in this equation $C \cos \left(v_{y} y\right) \neq 0$ so

$$
\begin{aligned}
-A v_{x} \sin \left(v_{x} a\right) & =0=\sin (0+q \pi), \quad q=0,1,2, \ldots \\
& \Rightarrow v_{x}=\frac{q \pi}{a}
\end{aligned}
$$

Above results yields
$\psi_{m}(\mathbf{r})=X(x) Y(y)=\left[A \cos \left(v_{x} x\right)\right]\left[C \cos \left(v_{y} y\right)\right]=A_{m}^{h} \cos \left(v_{x} x\right) \cos \left(v_{y} y\right)$
where $A_{m}^{h}$ is a normalization constant.
The eigenvalue is

$$
v_{m}^{2}=\pi^{2}\left(\frac{p^{2}}{a^{2}}+\frac{q^{2}}{b^{2}}\right) \equiv v_{p, q}^{2}
$$

where the parameters $p$ and $q$ are integers that $p+q \neq 0$ for $p, q=0,1,2, \ldots$.
To find normalization constant $A_{m}^{h}$ let's use normalization condition

$$
\begin{aligned}
& \frac{v_{m}^{2}}{S} \int_{S}\left|\psi_{m}(\mathbf{r})\right|^{2} d S=1 N \quad \text { where } \quad S=a . b \\
& \frac{v_{m}^{2}}{S} \int_{S}\left|\psi_{m}(\mathbf{r})\right|^{2} d S=\frac{v_{m}^{2}}{a \cdot b} \int_{S}\left(A_{m}^{h}\right)^{2} \cos ^{2}\left(v_{x} x\right) \cos ^{2}\left(v_{y} y\right) d x d y=1
\end{aligned}
$$

$$
\Rightarrow \frac{\left(A_{m}^{h}\right)^{2} \cdot v_{m}^{2}}{a \cdot b} \int_{x=0}^{x=a} \int_{y=0}^{y=b} \cos ^{2}\left(v_{x} x\right) \cos ^{2}\left(v_{y} y\right) d x d y=1
$$

$$
\Rightarrow \frac{\left(A_{m}^{h}\right)^{2} \cdot v_{m}^{2}}{a \cdot b} \int_{x=0}^{x=a} \cos ^{2}\left(v_{x} x\right) d x \int_{y=0}^{y=b} \cos ^{2}\left(v_{y} y\right) d y=1
$$

$$
\Rightarrow \frac{\left(A_{m}^{h}\right)^{2} \cdot v_{m}^{2}}{a \cdot b}\left[\left(\frac{x}{2}+\frac{1}{2} \frac{\sin \left(2 v_{x} x\right)}{2 v_{x} x}\right)_{x=0}^{x=a}\right]\left[\left(\frac{y}{2}+\frac{1}{2} \frac{\sin \left(2 v_{y} y\right)}{2 v_{y} y}\right)_{y=0}^{y=b}\right]=1
$$

$$
\begin{equation*}
\Rightarrow A_{m}^{h}=\frac{2}{v_{m}} \equiv A_{p, q}^{h} \tag{2.5}
\end{equation*}
$$

where the parameters $p$ and $q$ are integers that $p+q \neq 0$ for $p, q=0,1,2, \ldots$.
Take potential $\Psi_{m}(\mathbf{r})$ as $\Psi_{m}(\mathbf{r})=A_{m}^{h} \psi_{m}(\mathbf{r})$ where $A_{m}^{h}$ is a normalization constant. Then every components of TE time-domain modes can be written as
$\mathbf{E}_{z m}^{h}=0$
$v_{m}^{-1} \mathbf{E}_{m}^{h}=\left\langle-\partial_{\left(v_{m} c t\right)} h_{m}(z, t)\right\rangle\left[\sqrt[-2]{\varepsilon_{0}} \nabla_{\perp} \Psi_{m}(\mathbf{r}) \times \mathbf{z}\right]$
$v_{m}^{-1} \mathbf{H}_{m}^{h}=\left\langle\partial_{\left(v_{m}\right)} h_{m}(z, t)\right\rangle\left[\sqrt[-2]{\mu_{0}} \nabla_{\perp} \Psi_{m}(\mathbf{r})\right]$
$v_{m}^{-1} \mathbf{H}_{z m}^{h}=\left\langle h_{m}(z, t)\right\rangle\left[v_{m} \sqrt[-2]{\mu_{0}} \Psi_{m}(\mathbf{r})\right]$
where $\partial_{\left(v_{m} c t\right)}=\left(1 / c v_{m}\right) \partial_{t}, \partial_{\left(v_{m}\right)}=\left(1 / v_{m}\right) \partial_{z}$ and $c=\sqrt[-2]{\varepsilon_{0} \mu_{0}}$ is the speed of light in space. The potential $h_{m}(z, t)$ is obtained from Klein-Gordon equation
$\left(\partial_{v_{m} c t}^{2}-\partial_{v_{m} z}^{2}+1\right) h_{m}(z, t)=0$.

The problem (2.4) has a zero trivial solution. This solution corresponds to the eigenvalue $v_{0}^{2}=0$ and generates the problem, $\nabla_{\perp}^{2} \Psi_{0}(\mathbf{r})=0$ and $\left.\partial_{n} \Psi_{0}(\mathbf{r})\right|_{L}=0$, for a harmonic function $\Psi_{0}(\mathbf{r})$. he harmonic functions are attained as $\Psi_{0}(\mathbf{r})=c_{1}$ where $\mathbf{r} \in L+S$ and $c_{1}$ is a constant. The potential $\Psi_{0}(\mathbf{r})$ produce a $T E$ mode.
$\mathbf{E}_{0}^{h}(\mathbf{r}, z, t)=0, \quad \mathbf{H}_{0}^{h}(\mathbf{r}, z, t)=\mathbf{z} c_{1}$
where the modal amplitude of the field $\mathbf{H}_{0}^{h}$ is constant.
Weyl Theorem (Appendix C) means that TE time-domain modes are complete in Hilbert space $L_{2}$ (Weyl [6]).

### 2.4 TM Time-Domain Modes

Dirichlet boundary eigenvalue problem for Laplacian is expressed as

$$
\begin{align*}
& \left(\nabla_{\perp}^{2}+\kappa_{m}^{2}\right) \phi_{m}(\mathbf{r})=0 \\
& \left.\phi_{m}(\mathbf{r})\right|_{L}=0  \tag{2.9}\\
& \frac{\kappa_{m}^{2}}{S} \int_{S}\left|\phi_{m}(\mathbf{r})\right|^{2} d S=1 N
\end{align*}
$$

where $\kappa_{m}^{2}>0$ 's are eigenvalues for $m=1,2, \ldots$, the index $m$ is numerical values sorted by increasing order in the real axis and $\phi_{m}(\mathbf{r})$ 's are eigenvectors corresponding to the eigenvalues. The physical dimensions of field vector extentions $\mathbf{E}_{m}$ and $\mathbf{H}_{m}$ are $V_{m}^{-1}$ (Volt per meter) and $A_{m}^{-1}$ (Ampere per meter $)$, respectively. The solution $\phi_{0}(\mathbf{r})$ corresponding to eigenvalues $\kappa_{0}^{2}=0$ is zero.

Example 2.2 Consider previous example for the rectangular waveguide.
The solution to Dirichlet problem is same as Neumann's solution.
Since
$\nabla_{\perp}=\left(\frac{\partial}{\partial x}, \frac{\partial}{\partial y}\right)$
we can write Helmholtz equation as
$\left(\frac{\partial^{2}}{\partial x^{2}}+\frac{\partial^{2}}{\partial y^{2}}\right) \phi_{m}+\kappa_{m}^{2} \phi_{m}=0 \quad \Rightarrow \quad \frac{\partial^{2} \phi_{m}}{\partial x^{2}}+\frac{\partial^{2} \phi_{m}}{\partial y^{2}}+\kappa_{m}^{2} \phi_{m}=0$

By using the separation of variables method, choose $\phi_{m}(\mathbf{r})=X(x) Y(y)$ and substitute it into above equation yields
$\frac{\partial^{2}[X(x) Y(y)]}{\partial x^{2}}+\frac{\partial^{2}[X(x) Y(y)]}{\partial y^{2}}+\kappa_{m}^{2} X(x) Y(y)=0$
or equavalently
$Y \frac{\partial^{2} X}{\partial x^{2}}+X \frac{\partial^{2} Y}{\partial y^{2}}+\kappa_{m}^{2} X Y=0$.

Multiply both sides with $1 / X Y$ gives
$\underbrace{\frac{1}{X} \frac{\partial^{2} X}{\partial x^{2}}}_{-\kappa_{x}^{2}}+\underbrace{\frac{1}{Y} \frac{\partial^{2} Y}{\partial y^{2}}}_{-\kappa_{y}^{2}}+\kappa_{m}^{2}=0 \quad \Rightarrow \quad \kappa_{m}^{2}=\kappa_{x}^{2}+\kappa_{y}^{2}$

The eigenvalues $\kappa_{x}^{2}$ and $\kappa_{y}^{2}$ are choosen as negative because if we choose them as positive the solution of Neumann problem becomes $\phi_{m}(\mathbf{r})=\exp ( \pm i \kappa x) \exp ( \pm i \kappa y)$ which is a divergent function. So, it is not suitable for the EAE method.

Therefore, the solution of above equation can be written as
$X(x)=A \cos \left(\kappa_{x} x\right)+B \sin \left(\kappa_{x} x\right)$
$Y(y)=C \cos \left(\kappa_{y} y\right)+D \sin \left(\kappa_{y} y\right)$
or
$\phi_{m}(\mathbf{r})=X(x) Y(y)=\left[A \cos \left(\kappa_{x} x\right)+B \sin \left(\kappa_{x} x\right)\right]\left[C \cos \left(\kappa_{y} y\right)+D \sin \left(\kappa_{y} y\right)\right]$.

By using boundary condition, the constants of $\phi_{m}(\mathbf{r})$ can be found as stated below.
(i) $0 \leq x \leq a,\left.\phi_{m}(\mathbf{r})\right|_{y=0}=0$
$\left.\phi_{m}(\mathbf{r})\right|_{y=0}=\left.\left[A \cos \left(\kappa_{x} x\right)+B \sin \left(\kappa_{x} x\right)\right]\left[C \cos \left(\kappa_{y} y\right)+D \sin \left(\kappa_{y} y\right)\right]\right|_{y=0}=0$
which is equivalent to
$\left[A \cos \left(\kappa_{x} x\right)+B \sin \left(\kappa_{x} x\right)\right][C \cos (0)+D \sin (0)]=0$
or
$\left[A \cos \left(\kappa_{x} x\right)+B \sin \left(\kappa_{x} x\right)\right][C .1+D .0]=0$

In this equation $A \cos \left(\kappa_{x} x\right)+B \sin \left(\kappa_{x} x\right) \neq 0$ so
$C .1+D .0=0 \quad \Rightarrow \quad C=0$
(ii) $0 \leq x \leq a,\left.\phi_{m}(\mathbf{r})\right|_{y=b}=0$
$\left.\phi_{m}(\mathbf{r})\right|_{y=b}=\left.\left[A \cos \left(\kappa_{x} x\right)+B \sin \left(\kappa_{x} x\right)\right]\left[0 \cdot \cos \left(\kappa_{y} y\right)+D \sin \left(\kappa_{y} y\right)\right]\right|_{y=b}=0$
which is equivalent to
$\left[A \cos \left(\kappa_{x} x\right)+B \sin \left(\kappa_{x} x\right)\right]\left[D \sin \left(\kappa_{y} b\right)\right]=0$

Again in this equation $A \cos \left(\kappa_{x} x\right)+B \sin \left(\kappa_{x} x\right) \neq 0$ so $D \sin \left(\kappa_{y} b\right)=0=\sin (0+q \pi), \quad q=0,1,2, \ldots$

$$
\Rightarrow \kappa_{y}=\frac{q \pi}{a}
$$

(iii) $0 \leq y \leq b,\left.\phi_{m}(\mathbf{r})\right|_{x=0}=0$
$\left.\phi_{m}(\mathbf{r})\right|_{x=0}=\left.\left[A \cos \left(\kappa_{x} x\right)+B \sin \left(\kappa_{x} x\right)\right]\left[0 \cdot \cos \left(\kappa_{y} y\right)+D \sin \left(\kappa_{y} y\right)\right]\right|_{x=0}=0$
which is equivalent to
$[A \cos (0)+B \sin (0)]\left[D \sin \left(\kappa_{y} y\right)\right]=0$
or
$[A .1+B .0]\left[D \sin \left(\kappa_{y} y\right)\right]=0$

In this equation $D \sin \left(\kappa_{y} y\right) \neq 0$ so
$A .1+B .0=0 \quad \Rightarrow \quad A=0$
(iv) $0 \leq y \leq b,\left.\quad \phi_{m}(\mathbf{r})\right|_{x=a}=0$
$\left.\phi_{m}(\mathbf{r})\right|_{x=a}=\left.\left[0 \cdot \cos \left(\kappa_{x} x\right)+B \sin \left(\kappa_{x} x\right)\right]\left[0 \cdot \cos \left(\kappa_{y} y\right)+D \sin \left(\kappa_{y} y\right)\right]\right|_{x=a}=0$
which is equivalent to
$\left[B \sin \left(\kappa_{x} a\right)\right]\left[D \sin \left(\kappa_{y} y\right)\right]$

Again in this equation $D \sin \left(\kappa_{y} y\right) \neq 0$ so
$B \sin \left(\kappa_{x} a\right)=0=\sin (0+p \pi), \quad p=0,1,2, \ldots$

$$
\Rightarrow \kappa_{x}=\frac{p \pi}{a}
$$

Above results yields
$\phi_{m}(\mathbf{r})=X(x) Y(y)=\left[B \sin \left(\kappa_{x} x\right)\right]\left[D \sin \left(\kappa_{y} y\right)\right]=A_{m}^{e} \sin \left(\kappa_{x} x\right) \sin \left(\kappa_{y} y\right)$
where $A_{m}^{e}$ is a normalization constant.
Hence, the eigenvalue is found as
$\kappa_{m}^{2}=\pi^{2}\left(\frac{p^{2}}{a^{2}}+\frac{q^{2}}{b^{2}}\right) \equiv \kappa_{p, q}^{2}$
where the parameters $p$ and $q$ are integers that $p+q \neq 0$ for $p, q=0,1,2, \ldots$.
To find normalization constant $A_{m}^{e}$ let's use normalization condition

$$
\begin{align*}
\frac{\kappa_{m}^{2}}{S} \int_{S}\left|\phi_{m}(\mathbf{r})\right|^{2} d S & =1 N \quad \text { where } \quad S=a . b \\
\frac{\kappa_{m}^{2}}{S} \int_{s}\left|\phi_{m}(\mathbf{r})\right|^{2} d s & =\frac{\kappa_{m}^{2}}{a \cdot b} \int_{S}\left(A_{m}^{e}\right)^{2} \sin ^{2}\left(\kappa_{x} x\right) \sin ^{2}\left(\kappa_{y} y\right) d x d y=1 \\
& \Rightarrow \frac{\left(A_{m}^{e}\right)^{2} \cdot \kappa_{m}^{2}}{a \cdot b} \int_{x=0}^{x=a} \int_{y=0}^{y=b} \sin ^{2}\left(\kappa_{x} x\right) \sin ^{2}\left(\kappa_{y} y\right) d x d y=1 \\
& \Rightarrow \frac{\left(A_{m}^{e}\right)^{2} \cdot \kappa_{m}^{2}}{a \cdot b} \int_{x=0}^{x=a} \sin ^{2}\left(\kappa_{x} x\right) d x \int_{y=0}^{y=b} \sin ^{2}\left(\kappa_{y} y\right) d y=1 \\
& \Rightarrow \frac{\left(A_{m}^{e}\right)^{2} \cdot \kappa_{m}^{2}}{a \cdot b}\left[\left(\frac{x}{2}-\frac{1}{2} \frac{\sin \left(2 \kappa_{x} x\right)}{2 \kappa_{x} x}\right)_{x=0}^{x=a}\right]\left[\left(\frac{y}{2}-\frac{1}{2} \frac{\sin \left(2 \kappa_{y} y\right)}{2 \kappa_{y} y}\right)_{y=0}^{y=b}\right]=1 \\
& \Rightarrow A_{m}^{e}=\frac{2}{\kappa_{m}} \equiv A_{p, q}^{e} \tag{2.10}
\end{align*}
$$

where the parameters $p$ and $q$ are integers that $p+q \neq 0$ for $p, q=0,1,2, \ldots$. Then take the potential $\Phi_{m}(\mathbf{r})$ as $\Phi_{m}(\mathbf{r})=A_{m}^{e} \phi_{m}(\mathbf{r})$ where $A_{m}^{e}$ is normalization constant. Thus, every component of TM time-domain modes can be stated as
$\mathbf{H}_{z m}^{e}=0$
$\kappa_{m}^{-1} \mathbf{H}_{m}^{e}=\left\langle-\partial_{\left(\kappa_{m} c t\right)} e_{m}(z, t)\right\rangle\left[\mathbf{z} \times \sqrt[-2]{\mu_{0}} \nabla_{\perp} \Phi_{m}(\mathbf{r})\right]$
$\kappa_{m}^{-1} \mathbf{E}_{m}^{e}=\left\langle\partial_{\left(\kappa_{m} z\right)} e_{m}(z, t)\right\rangle\left[\sqrt[-2]{\varepsilon_{0}} \nabla_{\perp} \Phi_{m}(\mathbf{r})\right]$
$\kappa_{m}^{-1} \mathbf{E}_{m}^{e}=\left\langle e_{m}(z, t)\right\rangle\left[\kappa_{m} \sqrt[-2]{\varepsilon_{0}} \Phi_{m}(\mathbf{r})\right]$
where $\partial_{\left(\kappa_{m} c t\right)}=\left(1 / c \kappa_{m}\right) \partial_{t}, \quad \partial_{\kappa_{m} z}=\left(1 / \kappa_{m}\right) \partial_{z} \quad$ and $\quad c=\sqrt[-2]{\varepsilon_{0} \mu_{0}} \quad$ is the speed of light in space. The potential $e_{m}(z, t)$ is obtained from Klein-Gordon equation

$$
\begin{equation*}
\left(\partial_{v_{m} c t}^{2}-\partial_{v_{m} z}^{2}+1\right) e_{m}(z, t)=0 . \tag{2.12}
\end{equation*}
$$

Weyl Theorem (Appendix C) states that $T M$ time-domain modes are complete in Hilbert space $L_{2}$ (Weyl [6]).

## CHAPTER 3

## GENERAL PROPERTIES of TIME-DOMAIN MODES

### 3.1 Completeness of Time-Domain Modes

The set of TE and TM modes is complete due to the completeness of their generating potentials in the same energetic space. The completeness comes from Weyl theorem (Appendix C) in functional analysis about orthogonal detachments of Hilbert space $L_{2}$ Proof of completeness is stated in the references (Tretyakov [7]), (Aksoy and Tretyakov [8]), (Tretyakov[9]).
This energetic space can be specified by an inner product as
$\left(X_{1}, X_{2}\right)=\frac{1}{S} \int_{S}\left(\varepsilon_{0} \mathbf{E}_{1} \cdot \mathbf{E}_{2}^{*}+\mu_{0} \mathbf{H}_{1} \cdot \mathbf{H}_{2}^{*}\right) d s<\infty$
where $X_{1}=\operatorname{col}\left(\mathbf{E}_{1}, \mathbf{H}_{1}\right), X_{2}=\operatorname{col}\left(\mathbf{E}_{2}, \mathbf{H}_{2}\right)$ and the meaning of "col" is column and "." is scalar product of 3 -component vectors.

Since TE and Tm time-domain modes are complete, we can talk about their orthogonality. Let's choose the set of $T E$ and $T M$ modes (2.6) and (2.11) as
$\mathbf{G}=\left\{X_{m}^{h}\right\}_{m=0}^{\infty}, \quad \mathfrak{J}=\left\{X_{m}^{e}\right\}_{m=0}^{\infty}$.
respectively, where $X_{m}^{h}=\operatorname{col}\left(\mathbf{E}_{m}^{h}, \mathbf{H}_{m}^{h}\right)$ and $X_{m}^{e}=\operatorname{col}\left(\mathbf{E}_{m}^{e}, \mathbf{H}_{m}^{e}\right)$.
If we set $m=m^{\prime}$ and substitute the pair $X_{m}^{h}, X_{m^{\prime}}^{h}$ into the pair $X_{1}, X_{2}$ in equation (3.1), then it yields $\left(X_{m}^{h}, X_{m^{\prime}}^{h}\right)=0$. It shows that all elements of the set $\mathbf{G}$ are mutually orthogonal. The same is true for the elements of the set $\mathfrak{J}$. Let's take an element $X_{m}^{h}$ from the set $\mathbf{G}$ and take another element $X_{m}^{e}$ from the set $\mathfrak{I}$. Now, if we plug them into
equation (3.1), again we have the same result $\left(X_{m}^{h}, X_{m}^{e}\right)=0$ where $m=m^{\prime}$ is arbitrary.

Therefore, any pair of the TE and TM time-domain modes is orthogonal in the sense of inner product as in Figure 3.1.


Figure 3.1 Orthogonality of TE and TM modes
The proofs of the completeness of $\mathbf{G}$ and $\mathfrak{I}$ are in the studies of (Tretyakov [10]) and (Tretyakov [7]).

### 3.2 Energy Conservation Law for Time-Domain Modes

Let's select a volume $V$ located between two waveguide cross-sections along the coordinates $z$ and $z+\delta_{z}$ where $z$ and $\delta_{z}>0$ are arbitrary. After applying Poynting theorem (Appendix D) to Maxwell's equations for this volume with the limiting case of $\delta_{z} \rightarrow 0$, it yields law of conservation for any modal field energy as
$\frac{\partial}{\partial_{\xi}} P_{z}(\xi, \tau)+\frac{\partial}{\partial_{\tau}} W(\xi, \tau)=0$
where $P_{z}(\xi, \tau)$ and $W(\xi, \tau)$ is the $z$-component of Poynting vector averaged over the waveguide cross-section and averaged modal field energy stored in the same crosssection, respectively, and they defined as

$$
\begin{align*}
& W(\xi, \tau)=\left[\left(\partial_{\tau} f\right)^{2}+\left(\partial_{\xi} f\right)^{2}+f^{2}\right] / 2  \tag{3.4}\\
& P_{z}(\xi, \tau)=c\left(-\partial_{\tau} f\right)\left(\partial_{\xi} f\right) .
\end{align*}
$$

Equation (3.3) known as "continuity equation" holds for every modal field. Detailed solution is in (Akgun [2]) and (Eroglu [3]).

### 3.3 Initial Conditions for Klein-Gordon Equations

Klein-Gordon equation should has a pair of initial conditions like any other second order partial differential equations. Physically, they are involved in the excitation of a suitable signal source.Let's assume that this source is at rest before time $t=0$ but it starts excitation at time $t=0$. So, the initial condition can be written as
$\left.f(\xi, \tau)\right|_{\xi=0}=\left\{\begin{array}{lll}\varphi(\tau), & \tau \geq 0 & \text { for } \quad t \geq 0 \\ 0, & \tau<0 & \text { for } \quad t<0\end{array}\right.$
$\left.\frac{\partial}{\partial \tau} f(\xi, \tau)\right|_{\xi=0}=\left\{\begin{array}{lll}\hat{\varphi}(\tau), & \tau \geq 0 & \text { for } \quad t \geq 0 \\ 0, & \tau<0 & \text { for } \quad t<0\end{array}\right.$
where $\varphi(\tau)$ and $\hat{\varphi}(\tau)$ should be given and $\xi=0 \Rightarrow z=0$.

### 3.4 The Causality Principle

The solutions of Klein-Gordon equation is based on the causality principle. There are two kinds of causality. If the sources are zero at the beginning then all the fields are zero at weak causality. In our problem this correponds to $\tau<0$. The strong causality condition is the Einstein's postulate that claims any magnetic field propagates a signal in space with the speed of light $c$. In our problem, the source is in the waveguide crosssection at $\xi=0$. It states that the solution of Klein-Gordon must be zero at beyond the source point $\xi=0$, after $\xi=\tau$ (i.e., $z=c t$ ). Hence, if the signal propagates along Oz -axis then Klein-Gordon equation should be read as
$f(\xi, \tau)=\left\{\begin{array}{llr}f(\xi, \tau)=0 & \text { if } & \tau<0 \\ f(\xi, \tau) \neq 0 & \text { if } & 0 \leq \xi \leq \tau \\ f(\xi, \tau)=0 & \text { if } & \xi>\tau\end{array}\right.$

## CHAPTER 4

## MODAL AMPLITUDE PROBLEM

The wave equation which moves along the $z$-axis at a particular time $t$ is KleinGordon equation and it is defined by

$$
\left(\partial_{c t}^{2}-\partial_{z}^{2}+\kappa^{2}\right) F(z, t)=0
$$

where $c$ is the speed of light and $\kappa^{2}$ is an eigenvalue of Wave-Boundary Operators (which are inside the parenthesis of the equation).

Let us consider

$$
\begin{equation*}
\left(\partial_{c t}^{2}-\partial_{z}^{2}+1\right) F(z, t)=0 \tag{4.1}
\end{equation*}
$$

where $\kappa^{2} \equiv 1$. The function $F$ sought, in equation (4.1), is depend on two variables which are $z$ and $t: F \equiv F(z, t)$. To apply the method of separation of variables, we need to ask a question like "is it possible to write $(z, t)$ as a combination of $u(z, t)$ and $v(z, t) "$. To make it possible (i) we should take into new independent variables provided that (ii) these new variables can be separated in Klein-Gordon equation. In other words, we consider separation of variables in regard to the function $F$ as follows:

$$
\begin{equation*}
F \equiv F(z, t)=F[u(z, c t), v(z, c t)] \equiv F(u, v)=U(u) V(v) \tag{4.2}
\end{equation*}
$$

To answer above question, we should look at which form acquires equation (4.1) in terms of new variables, $u \equiv u(z, t)$ and $v \equiv v(z, t)$, as some functions of the original ones. For reforming equation (4.1) in terms of $u$ and $v$, we have to do the initial work as stated below:
$\frac{1}{c} \frac{\partial}{\partial t} F(u, v)=\frac{1}{c}\left\{\frac{\partial F}{\partial u} \frac{\partial u}{\partial t}+\frac{\partial F}{\partial v} \frac{\partial v}{\partial t}\right\}$

$$
\begin{aligned}
\frac{1}{c^{2}} \frac{\partial^{2}}{\partial t^{2}} F(u, v) & =\frac{1}{c^{2}}\left\{\frac{\partial^{2} F}{\partial u^{2}}\left(\frac{\partial u}{\partial t}\right)^{2}+\frac{\partial F}{\partial u} \frac{\partial^{2} u}{\partial t^{2}}+\frac{\partial^{2} F}{\partial v^{2}}\left(\frac{\partial v}{\partial t}\right)^{2}+\frac{\partial F}{\partial v} \frac{\partial^{2} v}{\partial t^{2}}\right\} \\
& +\frac{1}{c^{2}}\left\{\frac{\partial^{2} F}{\partial u \partial v} \frac{\partial u}{\partial t} \frac{\partial v}{\partial t}+\frac{\partial^{2} F}{\partial v \partial u} \frac{\partial v}{\partial t} \frac{\partial u}{\partial t}\right\}
\end{aligned}
$$

Since $u$ and $v$ are independent variables, the last equation can be rearrange as

$$
\begin{align*}
\frac{1}{c^{2}} \frac{\partial^{2}}{\partial t^{2}} F(u, v) & =\frac{1}{c^{2}}\left\{\frac{\partial^{2} F}{\partial u^{2}}\left(\frac{\partial u}{\partial t}\right)^{2}+\frac{\partial^{2} F}{\partial v^{2}}\left(\frac{\partial v}{\partial t}\right)^{2}+\frac{\partial F}{\partial u} \frac{\partial^{2} u}{\partial t^{2}}+\frac{\partial F}{\partial v} \frac{\partial^{2} v}{\partial t^{2}}\right. \\
& \left.+2 \frac{\partial^{2} F}{\partial u \partial v} \frac{\partial u}{\partial t} \frac{\partial v}{\partial t}\right\} \tag{4.3}
\end{align*}
$$

Equally,

$$
\begin{align*}
\frac{\partial^{2}}{\partial t^{2}} F(u, v)= & \left\{\frac{\partial^{2} F}{\partial u^{2}}\left(\frac{\partial u}{\partial t}\right)^{2}+\frac{\partial^{2} F}{\partial v^{2}}\left(\frac{\partial v}{\partial t}\right)^{2}+\frac{\partial F}{\partial u} \frac{\partial^{2} u}{\partial t^{2}}+\frac{\partial F}{\partial v} \frac{\partial^{2} v}{\partial t^{2}}\right. \\
& \left.+2 \frac{\partial^{2} F}{\partial u \partial v} \frac{\partial u}{\partial t} \frac{\partial v}{\partial t}\right\} \tag{4.4}
\end{align*}
$$

Substituting equations (4.3) and (4.4) in (4.1) lead us to

$$
\begin{array}{r}
{\left[\left(\frac{1}{c} \frac{\partial u}{\partial t}\right)^{2}-\left(\frac{\partial u}{\partial z}\right)^{2}\right] \frac{\partial^{2} F}{\partial u^{2}}+\left[\left(\frac{1}{c} \frac{\partial v}{\partial t}\right)^{2}-\left(\frac{\partial v}{\partial z}\right)^{2}\right] \frac{\partial^{2} F}{\partial v^{2}}+\left[\frac{1}{c^{2}} \frac{\partial^{2} u}{\partial t^{2}}-\frac{\partial^{2} u}{\partial z^{2}}\right] \frac{\partial F}{\partial u}} \\
+\left[\frac{1}{c^{2}} \frac{\partial^{2} v}{\partial t^{2}}-\frac{\partial^{2} v}{\partial z^{2}}\right] \frac{\partial F}{\partial v}+2\left[\frac{1}{c} \frac{\partial u}{\partial t} \frac{1}{c} \frac{\partial v}{\partial t}-\frac{\partial u}{\partial z} \frac{\partial v}{\partial z}\right] \frac{\partial^{2} F}{\partial u \partial v}+F=0 \tag{4.5}
\end{array}
$$

At this point, the following comments may be needed:

- Equation (4.5) is reformulated version of equation (4.1) with the transformation of the original variables to $u(z, t)$ and $v(z, t)$
- The terms in the square brackets of equation (4.5) has variable coefficients;
- The dependences of $u(z, t)$ and $v(z, t)$ should be defined to separate new variables $u$ and $v$ and find a solution of equation (4.5) in the form of
$F(u, v)=U(u) V(v)$. The definition can not be understood with a simple look at equation (4.5).

Klein-Gordon equation is well-studied about separation of variables in the mathematical book (Miller [11]). In this book, 11 orthogonal coordinate systems (Appendix A) were found which makes it possible to separate new variables, by using group theory approach. In addition, another set of nonorthogonal coordinate systems, in which separation of the variables is possible, can be attained with the aid of the same approach. The list of the orhogonal coordinate systems are cited below.

Group theory gives positive answer on this question. In the list, there is 11 versions of $c t$ and $z$ in terms of the functions of $u$ and $v$. Inverse functional depences, $u \equiv u(z, c t)$ and $v \equiv v(z, c t)$ can be found easily.

### 4.1 An Overview of Bessel Functions

### 4.1.1 Definition

Consider the homogenous linear $n-$ th order ordinary differential equation of is
$y^{(n)}+p_{n-1}(x) y^{(n-1)}+p_{n-2}(x) y^{(n-2)}+\cdots+p_{0}(x) y=0$.

In this case;

1. If the coefficients $p_{0}(x), p_{1}(x), \ldots, p_{n-1}(x)$ are not analytic at $x=x_{0}$ then $x_{0}$ is called a singular point of given differential equation.
2. If all the coefficients $p_{k}(x)$ are not analytic but all points $\left(x-x_{0}\right)^{n-k}$ for $k=$ $0,1, \ldots,(n-2)$ are analytic, then the point $x_{0}$ is regular singular.
3. If the point $x_{0}$ is neither an ordinary point nor a regular singular point of equation (4.6) then it is called as irregular singular point.

### 4.1.2 The Method of Frobenius

Let us consider a homogenous linear $2-$ nd order ordinary differential equation. If $x=$ 0 is a regular singular point then both $x P(x)$ and $x^{2} Q(x)$ can be written as power series $x P(x)=\sum_{n=0}^{\infty} P_{n} x^{n}, \quad|x|<r$
$x^{2} Q(x)=\sum_{n=0}^{\infty} Q_{n} x^{n}, \quad|x|<r$
or equivalenly,
$P(x)=\sum_{n=0}^{\infty} P_{n} x^{n-1}, \quad|x|<r, x \neq 0$
$Q(x)=\sum_{n=0}^{\infty} Q_{n} x^{n-2}, \quad|x|<r, \quad x \neq 0$.
Frobenius series solution of given differential equation is suggested as
$y(x)=x^{\alpha} \sum_{n=0}^{\infty} a_{n} x^{n+\alpha}=\sum_{n=0}^{\infty} a_{n} x^{n+\alpha}, \quad 0<x<r$
Differentiating $y$ with respect to $x$, in turn, yields
$y^{\prime}(x)=\sum_{n=0}^{\infty}(n+\alpha) a_{n} x^{n+\alpha-1}$,
$y^{\prime \prime}(x)=\sum_{n=0}^{\infty}(n+\alpha)(n+\alpha-1) a_{n} x^{n+\alpha-2}$
and substituting these results into the differential equation, we get

$$
\begin{align*}
& \sum_{n=0}^{\infty}(n+\alpha)(n+\alpha-1) a_{n} x^{n+\alpha-2}+\sum_{n=0}^{\infty} P_{n} x^{n-1} \sum_{n=0}^{\infty}(n+\alpha) a_{n} x^{n+\alpha-1} \\
& +\sum_{n=0}^{\infty} Q_{n} x^{n-2} \sum_{n=0}^{\infty} a_{n} x^{n+\alpha}=0 \tag{4.7}
\end{align*}
$$

In this equation, the second and third power series can be stated as
$\sum_{n=0}^{\infty} P_{n} x^{n-1} \sum_{n=0}^{\infty}(n+\alpha) a_{n} x^{n+\alpha-1}=\sum_{n=0}^{\infty} \sum_{m=0}^{n} P_{n-m} x^{n-m-1}(m+\alpha) a_{m} x^{m+\alpha-1}$

$$
\begin{equation*}
=\sum_{n=0}^{\infty}\left[\sum_{m=0}^{n}(m+\alpha) P_{n-m} a_{m}\right] x^{n+\alpha-2} \tag{4.8}
\end{equation*}
$$

and
$\sum_{n=0}^{\infty} Q_{n} x^{n-2} \sum_{n=0}^{\infty} a_{n} x^{n+\alpha}=\sum_{n=0}^{\infty}\left[\sum_{m=0}^{n} Q_{n-m} a_{m}\right] x^{n+\alpha-2}$,
respectively. Plugging equations (4.8) and (4.9) into (4.7) gives
$\sum_{n=0}^{\infty}\left\{(n+\alpha)(n+\alpha-1) a_{n}+\sum_{m=0}^{n}\left[(m+\alpha) P_{n-m}+Q_{n-m}\right] a_{m}\right\} x^{n+\alpha-2}=0$.

However, for equation (4.10) to hold the coefficient $x^{n+\alpha-2}, n=0,1, \ldots$ should be zero.
Our task is to find a recurrence relation involving $a_{n}$, for $n=0,1, \ldots$. So, setting $n=0$ equation (4.10) becomes
$\left[\alpha(\alpha-1)+\alpha P_{0}+Q_{0}\right] a_{0}=0$.
In this case, $a_{0}=0$ or $\alpha(\alpha-1)+\alpha P_{0}+Q_{0}=0$.
Setting $n=1$ in (4.10)
$(n+\alpha)(n+\alpha-1) a_{n}+\sum_{m=0}^{n}\left[(m+\alpha) P_{n-m}+Q_{n-m}\right] a_{m}=0$
or,
$a_{n}=-\frac{1}{(n+\alpha)(n+\alpha-1)+(n+\alpha) P_{0}+Q_{0}} \sum_{m=0}^{n-1}\left[(m+\alpha) P_{n-m}+Q_{n-m}\right] a_{m}$.

1. If $a_{0}=0$ then $a_{1}=a_{2}=\cdots=0$ and so $y(x)=0$.
2. If $a_{0} \neq 0$ then the indicial equation

$$
\begin{equation*}
\alpha(\alpha-1)+\alpha P_{0}+Q_{0}=0 \tag{4.11}
\end{equation*}
$$

is obtained.

We assume that $a_{0} \neq 0$, so equation (4.11) must be hold. Thus, $\alpha_{1}$ and $\alpha_{2}$ are the roots of the indicial equation associated with $x=0$.

### 4.1.3 Fuchs Theorem (Series Solution Near a Regular Singular Point)

We already know that if $x=0$ is a singular point for 2 nd order homogenous linear ordinary differential equation
$y^{\prime \prime}(x)+P(x) y^{\prime}(x)+Q(x) y(x)=0$,
then
$x P(x)=\sum_{n=0}^{\infty} P_{n} x^{n}, \quad x^{2} Q(x)=\sum_{n=0}^{\infty} Q_{n} x^{n}, \quad|x|<r$.

Let the indicial equation (4.11) has two real roots $\alpha_{1}$ and $\alpha_{2}$ where $\alpha_{1} \geq \alpha_{2}$.

1. If $\alpha_{1}-\alpha_{2}$ is not an integer, then there exist two linearly independent solutions of the form

$$
\begin{align*}
& y_{1}(x)=x^{\alpha_{1}} \sum_{n=0}^{\infty} a_{n} x^{n}, \quad a_{0} \neq 0,0<x<r  \tag{4.13}\\
& y_{2}(x)=x^{\alpha_{2}} \sum_{n=0}^{\infty} b_{n} x^{n}, \quad a_{0} \neq 0, \quad 0<x<r . \tag{4.14}
\end{align*}
$$

The coefficients $b_{n}$ is found by substituting $y_{1}(x)$ into the differential equation.
2. If $\alpha_{1}=\alpha_{2}=\alpha$ then the first and second linearly independent solutions are

$$
\begin{align*}
& y_{1}(x)=x^{\alpha_{1}} \sum_{n=0}^{\infty} a_{n} x^{n}, \quad a_{0} \neq 0, \quad 0<x<r \\
& y_{2}(x)=y_{1}(x) \ln x+x^{\alpha} \sum_{n=0}^{\infty} b_{n} x^{n}, \quad 0<x<r . \tag{4.15}
\end{align*}
$$

In this case, second solution $y_{2}(x)$ is not a Frobenius series solution.
3. If $\alpha_{1}-\alpha_{2}$ is an integer then

$$
\begin{align*}
& y_{1}(x)=x^{\alpha_{1}} \sum_{n=0}^{\infty} a_{n} x^{n}, \quad a_{0} \neq 0,0<x<r \\
& y_{2}(x)=a y_{1}(x) \ln x+x^{\alpha_{2}} \sum_{n=0}^{\infty} b_{n} x^{n}, \quad 0<x<r . \tag{4.16}
\end{align*}
$$

where $a$ is a constant that could be zero. The solution $y_{2}(x)$ can be a second solution of Frobenius series solution.

Consequently, the general solution of differential equation is

$$
\begin{equation*}
y(x)=C_{1}(x) y_{1}(x)+C_{2}(x) y_{2}(x) . \tag{4.17}
\end{equation*}
$$

### 4.1.4 Bessel Differential Equation and its Solution

The Bessel differential equation is the linear second-order ordinary differential equation given by
$x^{2} y^{\prime \prime}(x)+x y^{\prime}(x)+\left(x^{2}-p^{2}\right) y(x)=0, \quad x>0$
where $p \geq 0$ is a constant number. The method of series solution near a regular singular point is used for solving the Bessel differential equation.

Clearly,
$y^{\prime \prime}(x)+P(x) y^{\prime}(x)+Q(x) y(x)=0$
has a regular singular point at $x=0$ where $P(x)=\frac{1}{x}$ and $Q(x)=\frac{x^{2}-p^{2}}{x^{2}}$. Since,
$x P(x)=1=1+0 \cdot x+0 \cdot x^{2}+\cdots \quad \Rightarrow \quad P_{0}=1$
$x^{2} Q(x)=x^{2}-p^{2}=-p^{2}+0 \cdot x+x^{2}+0 \cdot x^{3}+\cdots \quad \Rightarrow \quad Q_{0}=-p^{2}$
$x P(x)$ and $x^{2} Q(x)$ are both analytic at the point $x=0$, can be written as a series expansion and these series are convergent for $|x|<\infty$, respectively.

By the indicial equation (4.11)
$\alpha(\alpha-1)+\alpha .1-p^{2}=0 \quad \Rightarrow \quad \alpha-p^{2}=0 \quad \Rightarrow \quad \alpha_{1}=p, \alpha_{2}=-p$.

The Bessel differential equation (4.18) has a series solution in the form of
$y_{1}(x)=x^{p} \sum_{n=0}^{\infty} a_{n} x^{n}=\sum_{n=0}^{\infty} a_{n} x^{n+p}, \quad a_{0} \neq 0, \quad 0<x<\infty$.

We begin by differentiating $y_{1}(x)$ in (4.23) with respect to $x$ to obtain
$y_{1}^{\prime}(x)=\sum_{n=0}^{\infty}(n+p) a_{n} x^{n+p-1}$
$y_{1}^{\prime \prime}(x)=\sum_{n=0}^{\infty}(n+p)(n+p-1) a_{n} x^{n+p-2}$.

Substituting $y_{1}(x), y_{1}^{\prime}(x)$ and $y_{1}^{\prime \prime}(x)$ into equation (4.18) yields

$$
\begin{aligned}
x^{2} \sum_{n=0}^{\infty}(n+p)(n+p-1) a_{n} x^{n+p-2} & +x \sum_{n=0}^{\infty}(n+p) a_{n} x^{n+p-1} \\
& +\left(x^{2}-p^{2}\right) \sum_{n=0}^{\infty} a_{n} x^{n+p}=0
\end{aligned}
$$

or more precisely,
$\sum_{n=0}^{\infty}\left[(n+p)(n+p-1)+(n+p)-p^{2}\right] a_{n} x^{n+p}+\sum_{n=0}^{\infty} a_{n} x^{n+p+2}=0$

To simplify the addition of two summations in (4.26), let's shift the index of second
$\sum_{n=0}^{\infty} a_{n} x^{n+p+2} \stackrel{n+2=m}{\Longrightarrow} \sum_{m=2}^{\infty} a_{n-2} x^{n+p}$.
With these changes equation (4.26) becomes
$x^{p}\left\{\sum_{n=0}^{\infty}\left[(n+p)(n+p-1)+(n+p)-p^{2}\right] a_{n} x^{n}+\sum_{n=2}^{\infty} a_{n-2} x^{n}\right\}=0$.

Since $x^{p} \neq 0$, then
$\sum_{n=0}^{\infty} n(n+2 p) a_{n} x^{n}+\sum_{n=2}^{\infty} a_{n-2} x^{n}=0$
holds for $x^{n} \neq 0, n=0,1, \ldots$.
Separating the terms corresponding to $n=0$ and $n=1$ and combining the rest under one summation, we have

$$
0 .(0+2 p) a_{0}+1 .(1+2 p) a_{1} x+\sum_{n=2}^{\infty}\left(n(n+2 p) a_{n}+a_{n-2}\right) x^{n}=0
$$

Setting the coefficients equal to zero gives
0. $(0+2 p) a_{0}=0$
where $a_{0} \neq 0$ is an arbitrary constant,

1. $(1+2 p) a_{1}=0 \quad \Rightarrow \quad a_{1}=0$
and the recurrence relation for $n \geq 2$ is
$n(n+2 p) a_{n}+a_{n-2}=0$
$a_{n}=-\frac{a_{n-2}}{n(n+2 p)}$

Solving the recurrence relation (4.29) for $a_{2 n+1}, n=0,1, \ldots$, we obtain $a_{2 n+1}=0$.

For $n=2$ and 4 . this gives
$a_{2}=-\frac{a_{0}}{2(2+2 p)}=-\frac{a_{0}}{2^{2} \cdot 1(1+p)}$
$a_{4}=-\frac{a_{2}}{4(4+2 p)}=-\frac{a_{2}}{2^{2} \cdot 2(2+p)}=(-1)^{2} \frac{a_{0}}{2^{4} \cdot 2!(1+p)(2+p)}$
it follows by induction
$a_{2 n}=(-1)^{n} \frac{a_{0}}{2^{2 n} \cdot n!(1+p)(2+p) \ldots(n+p)}$.

So, equation (4.23) is
$y_{1}(x)=a_{0} x^{p} \sum_{n=0}^{\infty}(-1)^{n} \frac{1}{n!(1+p)(2+p) \ldots(n+p)}\left(\frac{x}{2}\right)^{2 n}, \quad 0<x<\infty$.

In order to write simplified solution, the Gamma function can be used which is defined as
$\Gamma(p+1)=\int_{0}^{\infty} t^{p} e^{-t} d t, \quad p>0$
(Xie [12]). Using integration by part
$\Gamma(p+1)=-\left.t^{p} e^{-t}\right|_{t=0} ^{\infty}+\int_{0}^{\infty} p t^{p-1} e^{-t} d t=p \int_{0}^{\infty} t^{p-1} e^{-t} d t=p \Gamma(p)$,
is obtained. Thus
$\Gamma(n+p+1)=(n+p) \Gamma(n+p)=(n+p)(n+p-1) \Gamma(n+p-1)$

$$
=\cdots=(n+p)(n+p-1) \ldots(1+p) \Gamma(1+p) .
$$

In equation (4.31), for $p=0$
$\Gamma(1)=+\int_{0}^{\infty} e^{-t} d t=\left.e^{-t}\right|_{t=0} ^{\infty}=1$,
for $p=1$
$\Gamma(2)=2 . \Gamma(1)=1$,
for $p=2$
$\Gamma(3)=2 . \Gamma(2)=2.1=2!$,
for the case $p=k$ is an integer
$\Gamma(k+1)=k . \Gamma(k)=k!$.

Therefore, if we choose
$a_{0}=\left[2^{p} \Gamma(1+p)\right]^{-1}$
and, first Frobenius series solution is
$y_{1}(x)=J_{p}(x)$
where $J_{p}(x)$ is called as the Bessel function of the first kind and defined by
$J_{p}(x)=\frac{1}{2^{p} \Gamma(1+p)} x^{p} \sum_{n=0}^{\infty}(-1)^{n} \frac{1}{n!(1+p)(2+p) \ldots(n+p)}\left(\frac{x}{2}\right)^{2 n}$
or equivalently,
$J_{p}(x)=\sum_{n=0}^{\infty}(-1)^{n} \frac{1}{n!\Gamma(n+p+1)}\left(\frac{x}{2}\right)^{2 n+p}$
In Fuch's theory, linearly independent second solution changes according to the difference between the roots of the indicial equation (4.11). The difference $\alpha_{1}-\alpha_{2}=$ $2 p$ is not an integer, equal to zero or a positive integer

### 4.1.4.1 2p Nonbeing an Integer

Second Frobenius series solution (4.14) of the Bessel equation (4.18) is

$$
y_{2}(x)=x^{-p} \sum_{n=0}^{\infty} b_{n} x^{n}, \quad 0<x<\infty
$$

Similarly, in the previous section,
$b_{2 n-1}=0$,
and th recurrence relation is
$b_{2 n}=(-1)^{n} \frac{b_{0}}{2^{2 n} \cdot n!(1-p)(2-p) \ldots(n-p)}$.

Hence, the second Frobenius series solution is
$y_{2}(x)=b_{0} x^{-p} \sum_{n=0}^{\infty}(-1)^{n} \frac{1}{n!(1-p)(2-p) \ldots(n-p)}\left(\frac{x}{2}\right)^{2 n}, \quad 0<x<\infty$
or equivalently,
$y_{2}(x)=x^{-p} \sum_{n=0}^{\infty}(-1)^{n} \frac{1}{n!\Gamma(n-p+1)}\left(\frac{x}{2}\right)^{2 n-p}=J_{-p}(x), \quad 0<x<\infty$
where $b_{0}=\left[2^{-p} \Gamma(1-p)\right]^{-1}$.
As a result, the general solution of the Bessel equation is
$y(x)=C_{1} J_{p}(x)+C_{2} J_{-p}(x)$
or
$y(x)=D_{1} J_{p}(x)+D_{2} Y_{p}(x) y(x)=C_{1} J_{p}(x)+C_{2} J_{-p}(x)$
where $Y_{p}(x)$ is the Bessel function of the second kind of order two and represented by
$Y_{p}(x)=\frac{J_{p}(x) \cos (p \pi)-J_{-p}(x)}{\sin (p \pi)}$

### 4.1.4.2 $p=0, \alpha_{1}=\alpha_{2}=0$

The solution of first Frobenius series solution simplified from equation (4.32) as
$y_{1}(x)=J_{0}(x)=\sum_{n=0}^{\infty}(-1)^{n} \frac{1}{(n!)^{2}}\left(\frac{x}{2}\right)^{2 n}, \quad 0<x<\infty$
and linearly independent second solution is
$y_{2}(x)=y_{1}(x) \ln x+\sum_{n=0}^{\infty} b_{n} x^{n} . \quad 0<x<\infty$

We differentiate termwise to obtain
$y_{2}^{\prime}(x)=y_{1}^{\prime}(x) \ln x+\frac{y_{1}}{x}+\sum_{n=0}^{\infty} n b_{n} x^{n-1} \quad 0<x<\infty$
$y_{2}^{\prime \prime}(x)=y_{1}^{\prime \prime}(x) \ln x+\frac{2 y_{1}^{\prime}}{x}-\frac{y_{1}}{x^{2}}+\sum_{n=2}^{\infty} n(n-1) b_{n} x^{n-2}$.
and substituting them into the Bessel equation (4.18) for $p=0$ yields

$$
\begin{align*}
\left(x^{2} y_{1}^{\prime \prime}+x y_{1}^{\prime}\right. & \left.+x^{2} y_{1}\right) \ln x+2 x y_{1}^{\prime}+\sum_{n=2}^{\infty} n(n-1) b_{n} x^{n}+\sum_{n=1}^{\infty} n b_{n} x^{n} \\
& +\sum_{n=0}^{\infty} b_{n} x^{n+2}=0 \tag{4.41}
\end{align*}
$$

(Xie [12]).
Since $y_{1}(x)$ is a solution of the Bessel equation, we can write $x^{2} y_{1}^{\prime \prime}+x y_{1}^{\prime}+x^{2} y_{1}=0$ and

$$
\begin{equation*}
2 x y_{1}^{\prime}=2 x \sum_{n=0}^{\infty}(-1)^{n} \frac{1}{(n!)^{2}} \frac{2 n x^{2 n-1}}{2^{2 n}}=\sum_{n=1}^{\infty}(-1)^{n} \frac{4 n}{(n!)^{2}}\left(\frac{x}{2}\right)^{2 n} . \tag{4.42}
\end{equation*}
$$

Plugging above equations in equation (4.41) gives
$\sum_{n=1}^{\infty}(-1)^{n} \frac{4 n}{(n!)^{2}}\left(\frac{x}{2}\right)^{2 n}+\sum_{n=2}^{\infty} n(n-1) b_{n} x^{n}+\sum_{n=1}^{\infty} n b_{n} x^{n}+\sum_{n=2}^{\infty} b_{n-2} x^{n}=0$
where $x^{n}$ is need be equal zero to provide this equality for $n=0,1, \ldots$.
Next, if we separate the $x^{1}$ terms from others and set the coefficients of the power series equal to zeroe get
$b_{1}=0$,
$b_{2 n+1}=0$
and the recurrence relation
$(-1)^{n} \frac{4 n}{(n!)^{2}}\left(\frac{1}{2}\right)^{2 n}+[2 n(2 n-1)+2 n] b_{2 n}+b_{2 n-2}=0$
or
$b_{2 n}=(-1)^{n+1} \frac{1}{n(n!)^{2}}\left(\frac{1}{2}\right)^{2 n}-\frac{b_{2 n-2}}{(2 n)^{2}}$.
Assuming $b_{0}=0$ yields
$b_{2 n}=(-1)^{n+1} \frac{1+\frac{1}{2}+\frac{1}{3}+\cdots+\frac{1}{n}}{(n!)^{2}}\left(\frac{1}{2}\right)^{2 n}$.
Accordingly, the linearly independent second solution becomes
$y_{2}(x)=J_{0}(x) \ln x+\sum_{n=1}^{\infty}(-1)^{n+1} \frac{1+\frac{1}{2}+\frac{1}{3}+\cdots+\frac{1}{n}}{(n!)^{2}}\left(\frac{1}{2}\right)^{2 n} . \quad 0<x<\infty$
is a term of the Bessel funtion of the second kind of order zero. Or it can be expressed by
$y_{2}(x)=\frac{\pi}{2} Y_{0}(x)+(\ln 2-\gamma) J_{0}(x) \quad 0<x<\infty$
where $Y_{0}(x)$ is the Bessel funtion of the second kind of order zero. It is defined by $Y_{0}=\frac{2}{\pi}\left[\left(\ln \frac{x}{2}+\gamma\right) J_{0}(x)+\sum_{n=1}^{\infty}(-1)^{n+1} \frac{1+\frac{1}{2}+\frac{1}{3}+\cdots+\frac{1}{n}}{(n!)^{2}}\left(\frac{1}{2}\right)^{2 n}\right]$
where $\gamma=0,57721566490153 \ldots=\lim _{n \rightarrow \infty}\left(\sum_{k=1}^{n} \frac{1}{k}-\ln n\right)$ is Euler constant.
The general solution is
$y(x)=C_{0} J_{p}(x)+C_{2} Y_{0}(x)$.

### 4.1.4.3 $\boldsymbol{p}$ Being a Positive Integer

The solution of first Frobenius series solution was
$y_{1}(x)=J_{p}(x)=\sum_{n=0}^{\infty}(-1)^{n} \frac{1}{(n!)(n+p)!}\left(\frac{x}{2}\right)^{2 n+p} . \quad 0<x<\infty$

Linearly independent second solution is
$y_{2}(x)=a y_{1}(x) \ln x+x^{-p} \sum_{n=0}^{\infty} b_{n} x^{n} . \quad 0<x<\infty$

Our goal is to determine the coefficients $b_{n}$ by substituting $y_{2}(x)$ directly into equation (4.18). Let's take the derivatives of equation (4.48) with respect to $x$
$y_{2}^{\prime}(x)=a\left(y_{1}^{\prime}(x) \ln x+\frac{y_{1}}{x}\right)+\sum_{n=0}^{\infty}(n-p) b_{n} x^{n-p-1}$
$y_{2}^{\prime \prime}(x)=a\left(y_{1}^{\prime \prime}(x) \ln x+\frac{2 y_{1}^{\prime}}{x}-\frac{y_{1}}{x^{2}}\right)+\sum_{n=0}^{\infty}(n-p)(n-p-1) b_{n} x^{n-p-2}$
and substitute them into the Bessel equation (4.18)

$$
\begin{align*}
a\left(x^{2} y_{1}^{\prime \prime}+x y_{1}^{\prime}+\right. & \left.\left(x^{2}-p^{2}\right) y_{1}\right) \ln x+2 a x y_{1}^{\prime}+\sum_{n=0}^{\infty}(n-p)(n-p-1) b_{n} x^{n-p} \\
& +\sum_{\substack{n=\\
n=0}}^{\infty}(n-p) b_{n} x^{n-p}+\sum_{n=0}^{\infty} b_{n} x^{n-p-2}-\sum_{n=0}^{\infty} p^{2} b_{n} x^{n-p}=0 . \tag{4.51}
\end{align*}
$$

Since $y_{1}(x)$ is a solution of Bessel equation, we can write
$x^{2} y_{1}^{\prime \prime}+x y_{1}^{\prime}+\left(x^{2}-p^{2}\right) y_{1}=0$
and
$2 a x y_{1}^{\prime}=2 a x \sum_{n=0}^{\infty}(-1)^{n} \frac{(2 n+p) \cdot x^{2 n+p-1}}{n!(n+p)!2^{2 n+p}}=\sum_{n=1}^{\infty}(-1)^{n} \frac{2 a(2 n+p)}{n!(n+p)!}\left(\frac{x}{2}\right)^{2 n+p}$.

Multiply the equation by $x^{p}$ and substituting it and equation (4.52) into equation (4.51) gives
$\sum_{n=0}^{\infty}(-1)^{n} \frac{2^{p+2} a(2 n+p)}{n!(n+p)!}\left(\frac{x}{2}\right)^{2(n+p)}+\sum_{n=0}^{\infty} n(2 n-p) b_{n} x^{n}+\sum_{n=0}^{\infty} b_{n-2} x^{n}=0$
where $x^{n}$ is need to be equal zero to provide this equality for $n=0,1, \ldots$ (Xie [12]). In equation (4.53), there is no contribution of $x^{n}: 0 \leq n<2 p$.

Separating the terms correponding to $n=0$ and $n=1$ and setting the coefficients equal to zero yields
0. $(0-2 p) b_{0}=0$
where $b_{0}$ is an arbitrary and

1. $(1-2 p) b_{1}=0 \quad \Rightarrow \quad b_{1}=0$

Let's assume that $b_{0}=1$. The recurrence relation for $2 \leq n<2 p$ is
$n(n-2 p) b_{n}+b_{n-2}=0 \quad \Rightarrow \quad b_{n}=\frac{b_{n-2}}{n(2 p-n)}$.

Thus,
$b_{2}=\frac{b_{0}}{2(2 p-2)}=\frac{1}{2^{2} .1(p-1)} \quad(n=2)$,
$b_{4}=\frac{b_{2}}{4(2 p-4)}=\frac{1}{2^{4} \cdot 2!(p-1)(p-2)} \quad(n=4)$.

The pattern for the coefficients is now apparent. We find

$$
b_{2 k}=\frac{b_{2 k-2}}{2 k(2 p-2 k)}=\frac{1}{2^{2 k} \cdot k!(p-1)(p-2) \ldots(p-k)}=\frac{(p-k-1)!}{2^{2 k} \cdot k!(p-1)!}
$$

where $b_{2 n+1}=0$ is obvious for $n=0,1,2, \ldots$.
From the coefficient $x^{2 p}$ in equation (4.53)
$\frac{2^{p+1} a p}{p!}\left(\frac{x}{2}\right)^{2 p}+b_{2 p-2}=0$
then
$a=-2^{p-1}(p-1)!b_{2(p-1)}=-\frac{1}{2^{p-1}(p-1)}$.

Let's take the arbitrary contant $b_{2 p}=0$.
By the cofficient $x^{2(n+p)}, n \geq 1$
$(-1)^{n} \frac{2^{p+1} a(2 n+p)}{n!(n+p)!}\left(\frac{1}{2}\right)^{2(n+p)}+(2 n+2 p)(2 n) b_{2(n+p)}+b_{2(n-1+p)}=0$
Leave the term $b_{2(n+p)}$ alone
$b_{2(n+p)}=(-1)^{n+1} \frac{2^{p-1} a(2 n+p)}{n(n+p) n!(n+p)!}\left(\frac{1}{2}\right)^{2(n+p)}-\frac{b_{2(n-1+p)}}{2^{2} n(n+p)}$.
or equivalently,
$b_{2(n+p)}=(-1)^{n+1} \frac{2^{p-1} a A_{n}}{n!(n+p)!}\left(\frac{1}{2}\right)^{2(n+p)}$
where $A_{n}=\left(\frac{1}{1}+\frac{1}{2}+\cdots+\frac{1}{n}\right)\left(\frac{1}{1+p}+\frac{1}{2+p}+\cdots+\frac{1}{n+p}\right)$. So, linealy independent second solution is
$y_{2}(x)=a J_{p}(x) \ln x$

$$
\begin{equation*}
+x^{-p}\left(\sum_{n=0}^{p-1} \frac{(p-n-1)!}{n!(p-1)!}\left(\frac{x}{2}\right)^{2 n}+a \sum_{n=1}^{\infty}(-1)^{n+1} \frac{2^{p-1} a A_{n}}{n!(n+p)!}\left(\frac{x}{2}\right)^{2(n+p)}\right) \tag{4.56}
\end{equation*}
$$

or it can be expressed by

$$
\begin{align*}
y_{2}(x)= & a\left(J_{p}(x) \ln x+\frac{x^{-p}}{a} \sum_{n=0}^{p-1} \frac{(p-n-1)!}{n!(p-1)!}\left(\frac{x}{2}\right)^{2 n}\right. \\
& \left.\quad+\frac{1}{2} \sum_{n=0}^{\infty}(-1)^{n+1} \frac{\psi(n+1)+\psi(n+p+1)-\psi(p+1)+\gamma}{n!(n+p)!}\left(\frac{x}{2}\right)^{2 n+p}\right) \\
= & a\left(J_{p}(x) \ln \frac{x}{2}-\frac{1}{2} \sum_{n=0}^{p-1} \frac{(p-n-1)!}{n!}\left(\frac{x}{2}\right)^{2 n-p}\right. \\
& \quad-\frac{1}{2} \sum_{n=0}^{\infty}(-1)^{n} \frac{\psi(n+1)+\psi(n+p+1)}{n!(n+p)!}\left(\frac{x}{2}\right)^{2 n+p} \\
& \left.\quad-\frac{1}{2}(\gamma-\psi(p+1)-2 \ln 2) J_{p}(x)\right) \\
= & a\left(\frac{\pi}{2} Y_{p}(x)-\frac{1}{2}(\gamma-\psi(p+1)-2 \ln 2) J_{p}(x)\right), 0<x<\infty \tag{4.57}
\end{align*}
$$

where the function $\psi$ is a term of the Bessel function which is defined by
$\gamma+\psi(n+1)=\frac{1}{1}+\frac{1}{2}+\cdots+\frac{1}{n}$
$\psi(1)=-\gamma$
$\psi(n)=\frac{\Gamma^{\prime}(n)}{\Gamma(n)}$
and the Bessel function of the second kind is formulated as

$$
\begin{aligned}
& Y_{p}(x)=\frac{2}{\pi} J_{p}(x) \ln \frac{x}{2}-\frac{1}{\pi} \sum_{n=0}^{p-1} \frac{(p-n-1)!}{n!}\left(\frac{x}{2}\right)^{2 n-p} \\
&-\frac{1}{\pi} \sum_{n=0}^{\infty}(-1)^{n} \frac{\psi(n+1)+\psi(n+p+1)}{n!(n+p)!}\left(\frac{x}{2}\right)^{2 n+p} .
\end{aligned}
$$

The general solution can be expressed in terms of the Bessel function of the first and second kind as follows:
$y(x)=C_{1} J_{p}(x)+C_{2} Y_{p}(x)$.

This solution is the case of $\alpha_{1}-\alpha_{2}$ being an integer. The second solution contains the logaritmic function $\ln x$ (Xie [12]).
4.1.4.4 $p=k+\frac{1}{2}, k=0,1, \ldots$ and $\alpha_{1}-\alpha_{2}=2 k+1$ Being an Integer

The solution of first Frobenius series solution is
$y_{1}(x)=J_{k+\frac{1}{2}}(x)=\sum_{n=0}^{\infty}(-1)^{n} \frac{1}{(n!) \Gamma\left(n+k+\frac{3}{2}\right)}\left(\frac{x}{2}\right)^{2 n+k+\frac{1}{2}} \quad 0<x<\infty$
and first derivative of the solution is
$y_{1}^{\prime}(x)=\sum_{n=0}^{\infty}(-1)^{n} \frac{2 n+k+\frac{1}{2}}{2 n!\Gamma\left(n+k+\frac{3}{2}\right)}\left(\frac{x}{2}\right)^{2 n+k-\frac{1}{2}}$

By using equation (4.60), linearly independent second solution can be derived as
$y_{2}(x)=a y_{1}(x) \ln x+x^{-\left(k+\frac{1}{2}\right)} \sum_{n=0}^{\infty} b_{n} x^{n} . \quad 0<x<\infty$

Let's take the derivatives of this equation with respect to $x$
$y_{2}^{\prime}(x)=a\left(y_{1}^{\prime}(x) \ln x+\frac{y_{1}}{x}\right)+\sum_{n=0}^{\infty}\left(n-k-\frac{1}{2}\right) b_{n} x^{n-k-\frac{3}{2}}$
$y_{2}^{\prime \prime}(x)=a\left(y_{1}^{\prime \prime}(x) \ln x+\frac{2 y_{1}^{\prime}}{x}-\frac{y_{1}}{x^{2}}\right)+\sum_{n=0}^{\infty}\left(n-k-\frac{1}{2}\right)\left(n-k-\frac{3}{2}\right) b_{n} x^{n-k-\frac{5}{2}}$
and substitute them into the Bessel equation

$$
\begin{align*}
a\left(x^{2} y_{1}^{\prime \prime}+x y_{1}^{\prime}\right. & \left.+\left(x^{2}-p^{2}\right) y_{1}\right) \ln x+2 a x y_{1}^{\prime}+\sum_{n=0}^{\infty}(n-2 k-1) b_{n} x^{n-k-\frac{1}{2}} \\
& +\sum_{n=0}^{\infty} b_{n} x^{n-k+\frac{3}{2}}=0 . \tag{4.62}
\end{align*}
$$

Since $y_{1}(x)$ is a solution of Bessel equation, we can write

$$
x^{2} y_{1}^{\prime \prime}+x y_{1}^{\prime}+\left(x^{2}-p^{2}\right) y_{1}=0
$$

By multipling equation (4.62) by $x^{k+\frac{1}{2}}$, we obtain

$$
\begin{gather*}
a \sum_{n=0}^{\infty}(-1)^{n} \frac{2^{k+\frac{3}{2}}\left(2 n+k+\frac{1}{2}\right)}{n!\Gamma\left(n+k+\frac{3}{2}\right)!}\left(\frac{x}{2}\right)^{2 n+2 k+1}+\sum_{n=0}^{\infty} n(n-2 k-1) b_{n} x^{n} \\
+\sum_{n=0}^{\infty} b_{n} x^{n+2}=0 \tag{4.63}
\end{gather*}
$$

where $x^{2 n}=0=$ to hold this equality for $n=0,1, \ldots$. In equation (4.63), there is no contribution of $x^{n}: 0 \leq n<2 k+1$.

For $n=0$,
$x^{0}: 0 \cdot(0-2 k-1) b_{0}=0$
where $b_{0}$ is an arbitrary.
For $n=1$,
$x^{1}: 1 .(1-2 k-1) b_{1}=0 \quad \Rightarrow \quad b_{1}=0$

For $n$ : $2 \leq n<2 p$ we can easily see that

$$
b_{2 m+1}=-\frac{b_{2 m-1}}{(2 m+1)(2 m-2 k)}=0 . \quad 1<m<k
$$

From the coefficient $x^{2 k+1}$
$a \frac{k+\frac{1}{2}}{2^{k-\frac{1}{2}} \Gamma\left(k+\frac{3}{2}\right)}+(2 k+1) \cdot 0 \cdot b_{2 k+1}+b_{2 k-1}=0$
then
$a=0$

Plugging equation (4.64) into equation (4.63) and by the coefficient $x^{n}$
$n(n-2 p) b_{n}+b_{n-2}=0 \Rightarrow b_{2 m-1}=0$
and
$b_{2 m}=-\frac{b_{2 m-2}}{2^{2} m(m-p)} \quad$ for $m=1,2, \ldots$
or equivalently,
$b_{2 m}=(-1)^{m} \frac{b_{0}}{2^{2 m} m!(1-p)(2-p) \ldots(m-p)}$.
If we choose $b_{0}=\left(2^{-p} \Gamma(1-p)\right)^{-1}$, then
$b_{2 m}=\frac{(-1)^{m}}{m!\Gamma(m-p+1)}\left(\frac{1}{2}\right)^{2 m-p}$.

Thus, linealy independent second solution is
$y_{2}(x)=x^{-p} \sum_{n=0}^{\infty} b_{2 n} x^{-2 n}=\sum_{n=0}^{\infty} \frac{(-1)^{n}}{n!\Gamma(n-p+1)}\left(\frac{x}{2}\right)^{2 n+p}=J_{-p}(x) . \quad 0<x<\infty$

The general solution can be expressed by
$y(x)=C_{1} J_{p}(x)+C_{2} J_{-p}(x)$.
This solution is same as the case of $\alpha_{1}-\alpha_{2}$ being a positive integer. The second solution is a Frobenius seriesand does not contain the logaritmic function $\ln x$ (Xie [12]).

In application, the Bessel function is rarely in standart form.
Let's consider the second order ordinary differential equation
$\frac{d^{2} y}{d x^{2}}+\frac{1-2 \alpha}{x} \frac{d y}{d x}+\left(\beta \rho x^{\rho-1}+\frac{\alpha^{2}-p^{2} \rho^{2}}{x^{2}}\right) y=0, \quad x>0$
where $\alpha, \beta, p$ and $\rho$ are constant numbers.
Now, we will interchange the Bessel equation
$\xi^{2} \frac{d^{2} \eta}{d \xi^{2}}+\xi \frac{d \eta}{d \xi}+\left(\xi^{2}-p^{2}\right) \eta=0, \quad \xi>0$
to equation (4.65) as stated below.
By using change of variables method $\xi=\beta x^{\rho}$, we can write
$\xi \frac{d y}{d \xi}=\xi \frac{d y / d x}{d \xi / d x}=\beta x^{\rho} \frac{1}{\beta \rho x^{\rho-1}} \frac{d y}{d x}=\frac{x}{\rho} \frac{d y}{d x} \quad \Rightarrow \quad \xi \frac{d}{d \xi}=\frac{x}{\rho} \frac{d}{d x}$
which simplifies equation (4.66) as
$\xi \frac{d}{d \xi}\left(\xi \frac{d \eta}{d \xi}\right)+\left(\xi^{2}-p^{2}\right) \eta=0$
and then substitute equation (4.67) in eq. (4.68)
$\frac{x}{\rho} \frac{d}{d x}\left(\frac{x}{\rho} \frac{d \eta}{d x}\right)+\left(\beta^{2} x^{2 \rho}-p^{2}\right) \eta=0$.
For $\eta=x^{-\alpha} y$
$x \frac{d \eta}{d x}=x^{1-\alpha} \frac{d y}{d x}-\alpha x^{-\alpha} y$
and
$x \frac{d}{d x}\left(x \frac{d \eta}{d x}\right)=x^{2-\alpha} \frac{d^{2} y}{d x^{2}}+(1-2 \alpha) x^{1-\alpha} \frac{d y}{d x}+\alpha^{2} x^{-\alpha} y$.
Finally, the equation (4.66) becomes
$x^{2-\alpha} \frac{d^{2} y}{d x^{2}}+(1-2 \alpha) x^{1-\alpha} \frac{d y}{d x}+\alpha^{2} x^{-\alpha} y+\left(\beta^{2} x^{2 \rho}-p^{2}\right) \rho^{2} x^{-\alpha} y=0$.
which is equal to equation (4.65).

### 4.2 Expression of Modal Amplitudes via Bessel Functions

Let's investigate Miller's second case.
Case 2: $c t=u \cosh v, z=u \sinh v ; 0 \leq u<\infty,-\infty<v<\infty$. Initially, we need to find relations for $u(z, t)$ and $v(z, t)$. In order to do that, let us take square of two equations mentioned
$c^{2} t^{2}=u^{2} \cosh ^{2} v, \quad z^{2}=u^{2} \sinh ^{2} v$
and then
$c^{2} t^{2}-z^{2}=u^{2}\left(\cosh ^{2} v-\sinh ^{2} v\right)=u^{2}$.

Hence,
$u=\sqrt{c^{2} t^{2}-z^{2}}$
since $0 \leq u<\infty$, by definition. Similarly,
$\frac{z}{c t}=\frac{u \sinh v}{u \cosh v}=\tanh v$
and thus,
$v=\operatorname{arctanh} \frac{z}{c t} \equiv \frac{1}{2} \ln \frac{c t+z}{c t-z}$.

The next step is to calculate the partial derivatives of $u(z, t)$ and $v(z, t)$ in equation (4.5).
$\frac{1}{c} \frac{\partial u}{\partial t}=\frac{1}{c} \frac{\partial}{\partial t} \sqrt{c^{2} t^{2}-z^{2}}=\frac{c t}{\sqrt{c^{2} t^{2}-z^{2}}}=\frac{u \cosh v}{u}=\cosh v ;$
$\frac{1}{c} \frac{\partial v}{\partial t}=\frac{1}{c} \frac{\partial}{\partial t}\left(\frac{1}{2} \ln \frac{c t+z}{c t-z}\right)=-\frac{z}{c^{2} t^{2}-z^{2}}=-\frac{u \sinh v}{u^{2}}=-\frac{\sinh v}{u} ;$
$\frac{1}{c^{2}} \frac{\partial^{2} u}{\partial t^{2}}=\frac{1}{c} \frac{\partial}{\partial t}\left(\frac{1}{c} \frac{\partial u}{\partial t}\right)=\frac{1}{c} \frac{\partial}{\partial t} \cosh v=\frac{1}{c} \frac{\partial v}{\partial t} \sinh v=-\frac{\sinh ^{2} v}{u} ;$
$\frac{1}{c^{2}} \frac{\partial^{2} v}{\partial t^{2}}=\frac{1}{c} \frac{\partial}{\partial t}\left(\frac{1}{c} \frac{\partial v}{\partial t}\right)=\frac{1}{c} \frac{\partial}{\partial t}\left(-\frac{\sinh v}{u}\right)=\frac{\sinh 2 v}{u^{2}} ;$
$\frac{\partial u}{\partial z}=\frac{\partial}{\partial z} \sqrt{c^{2} t^{2}-z^{2}}=-\frac{z}{\sqrt{c^{2} t^{2}-z^{2}}}=-\frac{u \sinh v}{u}=-\sinh v ;$
$\frac{\partial v}{\partial z}=\frac{\partial}{\partial z}\left(\frac{1}{2} \ln \frac{c t+z}{c t-z}\right)=\frac{c t}{c^{2} t^{2}-z^{2}}=\frac{u \cosh v}{u^{2}}=\frac{\cosh v}{u} ;$
$\frac{\partial^{2} u}{\partial z^{2}}=-\frac{\cosh ^{2} v}{u} ;$
$\frac{\partial^{2} v}{\partial z^{2}}=\frac{\sinh 2 v}{u^{2}} ;$

Substituting these results in Klein-Gordon equation (4.5), a new representation in terms of the variables $u$ and $v$ arises as

$$
\begin{aligned}
& {\left[\cosh ^{2} v-\sinh ^{2} v\right] \frac{\partial^{2} F}{\partial u^{2}}+\left[\frac{\sinh ^{2} v}{u^{2}}-\frac{\cosh ^{2} v}{u^{2}}\right] \frac{\partial^{2} F}{\partial v^{2}}+\left[-\frac{\sinh ^{2} v}{u}+\frac{\cosh ^{2} v}{u}\right] \frac{\partial F}{\partial u}} \\
& \quad+\left[\frac{\sinh 2 v}{u^{2}}-\frac{\sinh 2 v}{u^{2}}\right] \frac{\partial F}{\partial v}+2\left[-\frac{\sinh v \cosh v}{u}+\frac{\sinh v \cosh v}{u}\right] \frac{\partial^{2} F}{\partial u \partial v} \\
& +F=0
\end{aligned}
$$

The last equation can be simplified as
$\frac{\partial^{2} F}{\partial u^{2}}+\frac{1}{u} \frac{\partial F}{\partial u}+F-\frac{1}{u^{2}} \frac{\partial^{2} F}{\partial v^{2}}=0$

Now, the solution of equation (4.78) can be obtained by separation of the variables ( $u, v$ ). So, if we choose
$F(u, v)=U(u) V(v)$,
then equation (4.78) can be rewritten as
$\left(\frac{\partial^{2}}{\partial u^{2}}+\frac{1}{u} \frac{\partial}{\partial u}+1-\frac{1}{u^{2}} \frac{V^{\prime \prime}(v)}{V(v)}\right) U(u)=0$.

In equation (4.80), $V^{\prime \prime}(v)=\alpha^{2} V(v)$ where $\alpha$ is a free parameter that is a constant of separation of the variables $u$ and $v$. Solution of this equation is
$V(v)=e^{ \pm \alpha v}$.

Thus, the function $U(u)$ satisfies Bessel equation

$$
\begin{equation*}
\left(\frac{\partial^{2}}{\partial u^{2}}+\frac{1}{u} \frac{\partial}{\partial u}+1-\frac{\alpha^{2}}{u^{2}}\right) U(u)=0 . \tag{4.82}
\end{equation*}
$$

which has two linearly independent solutions, in particularly,
$U(u)=A_{\alpha} J_{\alpha}(u)+B_{\alpha} Y_{\alpha}(u)$,
where $A_{\alpha}$ and $B_{\alpha}$ are arbitrary constants, $J_{\alpha}$ and $Y_{\alpha}$ are Bessel functions.

Example 4.1 Let us consider
$V(v)=e^{-\alpha v}, \quad U(u)=J_{\alpha}(\kappa u)$.

So
$F(z, t)=U(u) V(v)=e^{-\alpha v} J_{\alpha}(\kappa u)$.

Since $v=\frac{1}{2} \ln \frac{c t+z}{c t-z}, u=\sqrt{c^{2} t^{2}-z^{2}}$ and
$e^{-\alpha v}=e^{-\frac{\alpha}{2} \ln \frac{c t+z}{c t-z}}=e^{\ln \left(\frac{c t-z}{c t+z}\right)^{\alpha / 2}}=\left(\frac{c t-z}{c t+z}\right)^{\alpha / 2}$
then right-hand side of equation (4.84) become
$F(z, t)=\left(\frac{c t-z}{c t+z}\right)^{\alpha / 2} J_{\alpha}\left(\kappa \sqrt{c^{2} t^{2}-z^{2}}\right) \equiv F_{\alpha}(z, t)$
where $\alpha$ is free numerical parameter. The equation (4.85) transforms into
$F_{\alpha}(z, t)=J_{\alpha}(\kappa c t) \quad$ at $z=0$.

When we choose the free parameter $\alpha$ as integer: $\alpha=0,1,2, \ldots$, then a set of function

$$
\begin{equation*}
\left\{J_{n}(\kappa c t)\right\}_{n=0}^{\infty}, \tag{4.87}
\end{equation*}
$$

is complete. It means that arbitrary function of time $\varphi(t)$ can be expressed in the form of Neumann series, that is,
$\varphi(t)=\sum_{n=0}^{\infty} C_{n} J_{n}(\kappa c t)$
where $C_{n}$ 's are appropriate constants.

### 4.3 Differential Equation and Functions of Airy

Consider the Airy's differential equation which is named after the applied mathematician and astronomer G. B. Airy, who investigated the second order differential equation,
$y^{\prime \prime}-z y=0$.

Since the term $-z$ is analytic at $z=0$ then $z=0$ is an ordinary point for equation (4.88). Hence, we can express its general solution as a power series in the form
$y(z)=\sum_{n=0}^{\infty} a_{n} z^{n}$

The first and second derivatives of equation (4.89) with respect to z are
$y^{\prime}(z)=\sum_{n=1}^{\infty} n a_{n} z^{n-1}$
and
$y^{\prime \prime}(z)=\sum_{n=2}^{\infty} n(n-1) a_{n} z^{n-2}$,
respectively.
Substituting these power series into equation (4.88) we find
$\sum_{n=2}^{\infty} n(n-1) a_{n} z^{n-2}-z \sum_{n=0}^{\infty} a_{n} z^{n}=0$
or, equivalently,
$\sum_{n=2}^{\infty} n(n-1) a_{n} z^{n-2}-\sum_{n=0}^{\infty} a_{n} z^{n+1}=0$

The next goal is simplfy this expression by combining everything into a single summation. Let's shift the indeces so that the general term in each is a constant times $z^{n}$, that is,
$\sum_{n=0}^{\infty}(n+2)(n+1) a_{n+2} z^{n}-\sum_{n=1}^{\infty} a_{n-1} z^{n}=0$.

For the first summation, we split of the 0 -th term as
$\sum_{n=0}^{\infty}(n+2)(n+1) a_{n+2} z^{n}=2.1 \cdot a_{2}+\sum_{n=1}^{\infty}(n+2)(n+1) a_{n+2} z^{n}$

With these changes of indeces, equation (4.92) becomes
$2 a_{2}+\sum_{n=1}^{\infty}\left((n+2)(n+1) a_{n+2} z^{n}-a_{n-1} z^{n}\right)=0$
or
$2 a_{2}+\sum_{n=1}^{\infty}\left((n+2)(n+1) a_{n+2}-a_{n-1}\right) z^{n}=0$.

Since the power series is equal to zero, all coefficients are supposed to be equal to zero, too. Therefore,
$a_{2}=0$
and the recurrence relation is
$(n+2)(n+1) a_{n+2}=a_{n-1}$
or
$a_{n+2}=\frac{a_{n-1}}{(n+2)(n+1)}, \quad$ for all $n=1,2,3, \ldots$.

To find the series solution to equation (4.88), the values of $a_{n}$ 's are needed to determine.

For $n=1$, equation (4.93) is
$a_{3}=\frac{a_{0}}{3.2}$

Continuing,
$a_{4}=\frac{a_{1}}{4.3}$
$a_{5}=\frac{a_{2}}{5.4}=0$
$a_{6}=\frac{a_{3}}{6.5}=\frac{a_{0}}{(6.5)(3.2)}$
$a_{7}=\frac{a_{4}}{7.6}=\frac{a_{1}}{(7.6)(4.3)}$
$a_{8}=\frac{a_{5}}{8.7}=0$
$a_{9}=\frac{a_{6}}{9.8}=\frac{a_{0}}{(9.8)(6.5)(3.2)}$

The pattern for the coefficients is now apparent. So, let's write down all the cases:
(i) The terms $a_{2}, a_{5}, a_{8}, \ldots$ are equal to zero, that is,

$$
a_{3 n-1}=0 \text { for all } n=1,2, \ldots
$$

(ii) The terms $a_{3}, a_{6}, a_{9}, \ldots$ are multiples of $a_{0}$, so

$$
a_{3 n}=\frac{a_{0}}{[(3 n)(3 n-1)][(3 n-3)(3 n-4)] \ldots[6.5][3.2]}
$$

for all $n=1,2, \ldots$.
(iii)All the terms $a_{4}, a_{7}, a_{10}, \ldots$ are multiples of $a_{1}$ and so

$$
a_{3 n+1}=\frac{a_{1}}{[(3 n+1)(3 n)][(3 n-2)(3 n-3)] \ldots[7.6][4.3]}
$$

for all $n=1,2, \ldots$.
Finally, the general solution of Airy's equation is given by

$$
\begin{align*}
y(z)=a_{0}\left[\sum_{n=1}^{\infty} \frac{z^{3 n}}{(3 n)(3 n-1)(3 n-3)(3 n-4) \ldots 3.2}\right] \\
+a_{1}\left[\sum_{n=1}^{\infty} \frac{z^{3 n+1}}{(3 n+1)(3 n)(3 n-2)(3 n-3) \ldots 4.3}\right] . \tag{4.94}
\end{align*}
$$

and this is also called as the Airy functions.
After G. B. Airy, the Airy functions were well-studied by H.Jeffreys and J.C.P Miller. It was them who introduced the Airy functions $A i$ and $B i$ as special solutions of Airy's equation. So, the solution of the Airy's equation was expressed by
$y(z)=c_{1} A i+c_{2} B i, \quad z \in \mathbb{R}$
where $A i$ is the first solution of the Airy's differential equation which is called as the Airy function of first kind and $B i$ is the second solution of the differential equation which is named as the Airy function of second kind.

The Airy functions $A i$ and $B i$ have rather simple integral representations through sine, cosine and power functions as follows:
$A i(z)=\frac{1}{\pi} \int_{0}^{\infty} \cos \left(\frac{t^{3}}{3}+z t\right) d t, \quad z \in \mathbb{R}$
and

$$
\begin{equation*}
B i(z)=\frac{1}{\pi} \int_{0}^{\infty}\left[\sin \left(\frac{t^{3}}{3}+z t\right)+\exp \left(-\frac{t^{3}}{3}+z t\right)\right] d t, \quad z \in \mathbb{R} \tag{4.96}
\end{equation*}
$$

### 4.4 Expression of Modal Amplitudes via Airy Functions

Let's investigate Miller's fifth case.
Case 5: $c t=\frac{1}{2}(u-v)^{2}+u+v, \quad z=-\frac{1}{2}(u-v)^{2}+u+v,-\infty<u, v<\infty$. First, the relation for $u(z, t)$ and $v(z, t)$ should be found. So,
$c t+z=2(u+v), \quad c t-z=(u-v)^{2}$

Adding above equations side by side gives
$u=\frac{c t+z}{4}-\frac{\sqrt{c t-z}}{2}$
and
$v=\frac{c t+z}{4}+\frac{\sqrt{c t-z}}{2}$

Taking partial derivatives of equation (4.97) and (4.98) to substitute them into equation (4.5)

$$
\begin{aligned}
& \frac{1}{c} \frac{\partial u}{\partial t}=\frac{1}{4}-\frac{1}{4 \sqrt{c t-z}}=\frac{1}{4}-\frac{1}{4(u-v)} \\
& \frac{1}{c} \frac{\partial v}{\partial t}=\frac{1}{4}+\frac{1}{4 \sqrt{c t-z}}=\frac{1}{4}+\frac{1}{4(u-v)} \\
& \frac{1}{c^{2}} \frac{\partial^{2} u}{\partial t^{2}}=\frac{1}{c} \frac{\partial}{\partial t}\left(\frac{1}{c} \frac{\partial u}{\partial t}\right)=\frac{1}{32(u-v)^{3 / 2}} \\
& \frac{1}{c^{2}} \frac{\partial^{2} v}{\partial t^{2}}=\frac{1}{c} \frac{\partial}{\partial t}\left(\frac{1}{c} \frac{\partial v}{\partial t}\right)=-\frac{1}{32(u-v)^{3 / 2}} \\
& \frac{\partial u}{\partial z}=\frac{1}{4}+\frac{1}{4(u-v)} \\
& \frac{\partial v}{\partial z}=\frac{1}{4}-\frac{1}{4(u-v)}
\end{aligned}
$$

$\frac{\partial^{2} u}{\partial z^{2}}=\frac{1}{32(u-v)^{3 / 2}}$
$\frac{\partial^{2} v}{\partial z^{2}}=-\frac{1}{32(u-v)^{3 / 2}}$
Substituting above results to equation (4.5) yields
$\frac{1}{4(u-v)} \frac{\partial^{2} F}{\partial u^{2}}-\frac{1}{4(u-v)} \frac{\partial^{2} F}{\partial v^{2}}+F=0$
(Aksoy and Tretyakov [8]), (Akgun [2]).
Now, the solution of equation (4.99) can be obtained by separation of the variables $(u, v)$. So, if we choose
$F(u, v)=U(u) V(v)$
then equation (4.99) can be rewritten as
$\frac{1}{U(u)} \frac{d^{2} U(u)}{d u^{2}}+4 u=\frac{1}{V(v)} \frac{d^{2} V(v)}{d v^{2}}+4 v=4 \alpha$.
where $\alpha$ is a constant of separation of the variables.
At this point, it is reasonable to change notations for the variables $u$ and v as
$\bar{u}=\sqrt[3]{4}(\alpha-u) \quad$ and $\quad \bar{v}=\sqrt[3]{4}(\alpha-v)$,
then equation (4.100) becomes
$\frac{d^{2} U(\bar{u})}{d \bar{u}^{2}}-\bar{u} U(\bar{u})=0 \quad$ and $\quad \frac{d^{2} V(\bar{v})}{d \bar{v}^{2}}-\bar{v} V(\bar{v})=0$.

Since equation (4.101) is standart Airy's differential equation, it has two linearly independet solution in the form of
$U(\bar{u})=C_{1} A i(\bar{u})+D_{1} B i(\bar{u})$
and

$$
\begin{equation*}
V(\bar{v})=C_{2} A i(\bar{v})+D_{2} B i(\bar{v}) \tag{4.103}
\end{equation*}
$$

where $C_{1}, C_{2}, D_{1}$ and $D_{2}$ are arbitrary constants and $A i$ and $B i$ are Airy Functions. Consequently, we shown that Miller's fifth case can be written in terms of Airy Functions.

## CHAPTER 5

## GRAPHICS

The energy $W_{\alpha}$ and surplus of energy $s W_{\alpha}$ are exhibited in the following figures for the changing values of $\alpha$. The Mapple programme (Appendix B) is used for drawings.


Figure 5.1 Time dependence of $W_{3}$ and $s W_{3}$ for $0 \leq \tau \leq 15$.


Figure 5.2 Time dependence of $W_{3}$ and $s W_{3}$ for $0 \leq \tau \leq 50$.


Figure 5.3 Time dependence of $W_{4}$ and $s W_{4}$ for $0 \leq \tau \leq 15$.


Figure 5.4 Time dependence of $W_{4}$ and $s W_{4}$ for $0 \leq \tau \leq 50$.

The following figures are exhibited the changes of modal ampitudes $A$ and $B$ for $\alpha=$ $0,1,2, \ldots$ which are depend on the solution $F$ of KGE functions, in the cylindrical and spherical form. These oscillations are in time $t$ (i.e., in $\tau$ ) and axial coordinate $z$ (i.e., in $\xi)$.


Figure 5.5 The changes of modal amplitudes in interval for $0 \leq \tau \leq 15$ where $\alpha=3$.


Figure 5.6 The changes of modal amplitudes in interval for $0 \leq \tau \leq 30$ where $\alpha=3$.


Figure 5.7 The changes of modal amplitudes in interval for $10 \leq \tau \leq 20$ where $\alpha=0$, $\xi=10$.


Figure 5.8 The changes of modal amplitudes in interval for $10 \leq \tau \leq 30$ where $\alpha=0$, $\xi=10$.

In the following graphs, the speed $v_{\alpha}$ is calculated with transverse $w_{\alpha}^{t}$ and longitudinal $w_{\alpha}^{l}$ parts of energy, separately.


Figure 5.9 The change of $v_{0}, w_{0}^{t}$ and $w_{0}^{l}$ in $0 \leq \xi \leq \tau$ where $\tau=10$ and $\alpha=0$


Figure 5.10 The change of $v_{1}, w_{1}^{t}$ and $w_{1}^{l}$ in $0 \leq \xi \leq \tau$ where $\tau=10$ and $\alpha=1$

## CHAPTER 6

## RESULTS AND DISCUSSION

In this study, the time-domain waveguide modes are expressed analytically by a method of Evolutionary Approach to Electromagnetics (EAE). A hollow waveguide is considered with the perfect electric conductor surfaces.

Initialy, relevant definition of the waveguide were given. To obtain the components of transverse electric (TE) and transverse magnetic (TM) time-domain modes, Maxwell's equation and fields are decomposed into their transverse and longitudinal parts. Two main problems are introduced to analyze transferring of transient signals along hollow waveguide. Modal bases are determined by obtaining the solution of Dirichlet and Neumann boundary eigenvalue problems. Solving Klein-Gordon Equation yields modal amplitudes as a function of time and axial coordinate . As a result, every component of time-domain modes are expressed as a product of modal bases and modal amplitudes.

According to this new method, we expressed the solutions of Klein-Gordon equations (modal amplitudes) in terms of some sellected proper functions. By using Miller's list, the modal amplitudes were obtained in terms of Airy and Bessel functions with especially sellected parameters. The informations such as type (or mode) of wave spanning over the waveguide or form and dimension of waveguide cross-section can be attained by using modal amplitudes.
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## APPENDIX-A

## MILLER'S LIST

## Coordinate Systems for Separation of Variables in Klein-Gordon Equation

1. $c t=u, z=v: U(u) V(v)$ is a product of the exponentials
2. $\quad c t=u \cosh v, \quad z=u \sinh v ; \quad 0 \leq u<\infty, \quad-\infty<v<\infty: \quad U(u) V(v)$ is a product of Bessel function and an exponential.
3. $c t=\frac{1}{2}\left(u^{2}+v^{2}\right), z=u v ;-\infty<u<\infty, 0 \leq v<\infty: U(u) V(v)$ is a product of parabolic cylinder functions.
4. $c t=u v, z=\frac{1}{2}\left(u^{2}+v^{2}\right) ;-\infty<u<\infty, 0 \leq v<\infty: U(u) V(v)$ is a product of parabolic cylinder functions.
5. $c t=\frac{1}{2}(u-v)^{2}+u+v, \quad z=-\frac{1}{2}(u-v)^{2}+u+v ; \quad-\infty<u, v<\infty$ : $U(u) V(v)$ is a product of Airy functions.
6. $2 c t=\cosh \frac{u-v}{2}+\sinh \frac{u+v}{2}, \quad 2 z=\cosh \frac{u-v}{2} \sinh \frac{u+v}{2}: \quad-\infty<u, v<\infty$ : $U(u) V(v)$ is a product of Mathieu functions.
7. $c t=\sinh (u-v)+\frac{1}{2} e^{u+v}, \quad z=\sinh (u-v)-\frac{1}{2} e^{u+v}: \quad-\infty<u, v<\infty$ : $U(u) V(v)$ is a product of Bessel functions.
8. $c t=\cosh (u-v)+\frac{1}{2} e^{u+v}, \quad z=\cosh (u-v)-\frac{1}{2} e^{u+v}: \quad-\infty<u, v<\infty$ : $U(u) V(v)$ is a product of Bessel functions
9. $c t=\sinh u \cosh v, z=\cosh u \sinh v ;-\infty<u, v<\infty: U(u) V(v)$ is a product of Mathieu functions.
10. $c t=\cosh u \cosh v, z=\sinh u \sinh v ;-\infty<u<\infty, 0 \leq v<\infty: U(u) V(v)$ is a product of Mathieu functions.
11. $c t=\cos u \cos v, \quad z=\sin u \sin v ; \quad 0<u<2 \pi, \quad 0 \leq v<\pi: U(u) V(v)$ is a product of Mathieu functions.

## APPENDIX-B

## MAPPLE CODES

Mapple Code for Figure 5.1 and Figure 5.2

```
> ##f22:Bi(u)*Bi(v)
> restart;
    alpha:=-1;
    y1:=-0.15;
    y2:=0.34;
    tau1:=0;
    tau2:=15;
    u:=(4^(1/3))*(alpha-(tau+xi)/4+(((tau-xi)^(1/2))/2));
    v:=(4^(1/3))*(alpha-(tau+xi)/4-(((tau-xi)^(1/2))/2));
    f:=AiryBi((4^(1/3))*(alpha-(tau+xi)/4+(()
xi)^(1/2))/2)))*AiryBi ((4^(1/3))*(alpha-(tau+xi)/4-(((tau-
xi)^(1/2))/2)));
    A:=-diff(f,tau);
    B:=diff(f,xi);
    sW4:=(1/2)* ((A)^2-(B)^2);
    WW:=(1/2)* (f^2+A^2+B^2);
    N:=1500; #N=numpoints
    th:=2; #th is thickness
    h4:=piecewise(0>tau,0,f);
    xi:=tau-0.05;
```

with (plots) :
\#a4:=plot(h4,tau=tau1..tau2, numpoints=N, thickness=th, view=[ tau1..tau2, y1..y2], axes=boxed, legend=typeset (F4) , color=blac k , symbol=box, linestyle=solid) ;
\#t7:=plot(ttau, tau=tau1..tau2, numpoints=N, thickness=th, view $=\left[\right.$ tau1..tau2, y1 . $\mathrm{y}^{2}$ ], axes=boxed, legend=typeset (A4) , color=bl ack,symbol=box,linestyle=dot);
\#t8:=plot(txi, tau=tau1..tau2, numpoints=N, thickness=th, view= [tau1..tau2, y1. . y2], axes=boxed, legend=typeset (B4) , color=bla ck, symbol=box,linestyle=dash);
w: =plot (WW, tau=tau1..tau2, numpoints=N, thickness=th, view=[ta u1 . .tau2, y1..y2], axes=boxed, legend=typeset (W[4]) , color=blac k , symbol=box, linestyle=dash) ;

```
SW:=plot(sW4, tau=tau1..tau2,numpoints=N, thickness=th , view=[
tau1..tau2,y1..y2],axes=boxed,legend=typeset(sW[4]) , color=b
lack,symbol=box,linestyle=solid);
```

display (\{SW,w\});

Mapple Code for Figure 5.3 and Figure 5.4

```
> ##f21:Bi(u)*Ai(v)
> restart;
alpha:=-1;
y1:=-0.64;
y2:=0.16 ;
tau1:=0;
tau2:=15;
u:=(4^(1/3))*(alpha-(tau+xi) / 4+(((tau-xi)^(1/2))/2));
v:=(4^(1/3))*(alpha-(tau+xi)/4-(((tau-xi)^(1/2))/2));
f3:=AiryBi ((4^(1/3))* (alpha-(tau+xi) /4+(((tau-
xi)^(1/2)) /2)) )*AiryAi ((4^(1/3))* (alpha- (tau+xi) /4-(((tau-
xi)^(1/2))/2)));
ttau:=-diff(f3,tau);
txi:=diff(f3,xi);
```

```
N:=1500; #N=numpoints
th:=2; #th is thickness
h3:=piecewise(0>tau,0,f3);
xi:=tau-0.05;
```

with (plots) :
a3: =plot (h3, tau=tau1. .tau2, numpoints $=\mathrm{N}$, thickness=th, view=[t
au1..tau2, y1..y2], axes=boxed,legend=typeset (F3) , color=black
,symbol=box,linestyle=solid);
t5:=plot(ttau, tau=tau1..tau2, numpoints=N, thickness=th, view=
[tau1..tau2,y1..y2], axes=boxed, legend=typeset (A3) , color=bla
ck,symbol=box,linestyle=dot);
t6:=plot(txi,tau=tau1..tau2, numpoints=N, thickness=th, view=[
tau1..tau2, y1. y2], axes=boxed, legend=typeset (B3) , color=blac
k,symbol=box,linestyle=dash) ;
display(\{a3,t5,t6\});

## APPENDIX-C

## WEYL THEOREM

Let $G$ be a Lie group and let $\rho$ be a group representation of $\boldsymbol{G}$ on $\mathbb{C}^{n}$ (for some natural number $n$ ) which is continuous in the sense that the function $G \times \mathbb{C}^{n} \rightarrow \mathbb{C}^{n}$ defined by $(g, v) \rightarrow \rho(g)(v)$ is continuous. Then for each $v \in \mathbb{C}^{n}$ and each $\alpha \in\left(\mathbb{C}^{n}\right)^{*}$, the function $G \rightarrow \mathbb{C}^{n}$ defined by $g \rightarrow \alpha(\rho(g)(v))$ is continuous. The vector space span of all such functions is called the space of representative functions.

The Weyl theorem says that, if $G$ is compact, then

1) The representative functions are dense in the space of all continuous functions, with respect to the supremum norm;
2) The representative functions are dense in the space $L^{2}(G)$ of all squareintegrable functions, with respect to aHaar measure on $G$;
3) The vector space span of the characters of the irreducible continuous representations of $G$ are dense in the space of all continuous functions from $G$ into $\mathbb{C}^{n}$ which are constant on each conjugacy class of $G$, with respect to the supremum norm.

This theorem is easy to deduce from the Stone-Weierstrass theorem if it is assumed that $G$ is a matrix group. On the other hand, it is a corollary of the Weyl theorem that every compact Lie group is isomorphic to some matrix group

## APPENDIX-D

## POYNTING'S THEOREM (ENERGY CONSERVATION THEOREM)

Poynting's theorem is a powerful statement of energy conservation. It can be used to relate power absorption in an object to incident field. The theorem allows us to calculate the rate at which energy changes within a given volume due to the action of electromagnetic fields.. The theorem says if $S$ is any closed mathematical surface and $V$ is the volume inside $S$, then
$\frac{\partial}{\partial t} \int_{V}\left(W_{c}+\varepsilon_{0} \mathbf{E} . \mathbf{E}+\mu_{0} \mathbf{H} . \mathbf{H}\right) d V+\oint_{S}(\mathbf{E} \times \mathbf{H}) d S=0$
where $W_{c}$ is the energy possessed by charged particles at a given point in $\mathrm{V}, \varepsilon_{0} \mathbf{E}$. $\mathbf{E}$ is the enery stored in the $\mathbf{E}$ electric field at a given point in $V$ and $\mu_{0} \mathbf{H} . \mathbf{H}$ is the enery stored in the $\mathbf{H}$ emagnetic field at a given point in $V$. The first integral gives the rate at which energy stored within the fields changes within the volume $V$ and the second integral gives the rate at which energy flows across the surface $S$ enclosing $V$.

PERSONAL INFORMATION
Name Surname : Sedef ŞEMŞíT

Date of birth and place : 29/09/1990 - Uşak
Foreign Languages : English
E-mail : sedefsemsit @ gmail.com

## EDUCATION

| Degree | Department | University | Date of <br> Graduation |
| :--- | :--- | :--- | :--- |
| Undergraduate | Mathematics | Fatih University | 2013 |

## PUBLISHMENTS

## Conference Papers

1. 

S. Semsit, E. Eroglu, K. Koklu, O. Isik, (2014) 3rd International Eurasian Conference on Mathematical Sciences and Applications (IECMSA-2014 Book of Abstracts), "Wave Boundary Operators for Time-Domain Electromagnetic Fields", 259 pp., Vienna, Austria, 2528 August,.
3.
S. Semsit, E. Eroglu, K. Koklu, O. Isık, (2014) 3rd International Eurasian Conference on Mathematical Sciences and Applications (IECMSA-2014 Book of Abstracts), "Decomposition of Components of Energy During Signal Transferring", 260 pp., Vienna, Austria, 25-28 August.
E. Eroglu, S. Semsit, E. Sener, U.S. Sener, (2016) World Academy of Science, Engineering and Technology - International Journal of Mathematical and Computational Sciences, "Transferring of Transient Signals Along Hollow Waveguide", 3(5)

